

**NLC - The Next Linear Collider Project**



# Stability and Ground Motion Challenges in Linear Colliders

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SLAC

for the NLC collaboration

ICFA Nanobeam 02

Lausanne, September 2, 2002



# Contents

- Brief review of natural ground motion and vibrations and their influence on LC
  - Effects of fast motion
  - R&D aimed to ensure NLC stability
  - Particular case of Final Doublet (FD)

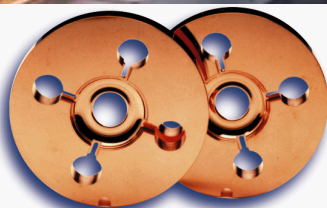
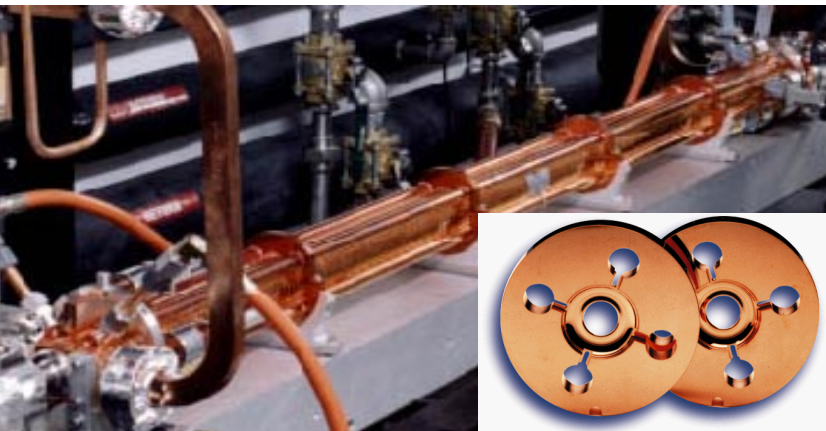


# Linear Colliders - two main challenges

- **Energy** - need to reach at least 500 GeV CM (as a start)
- **Luminosity** - need to reach  $10^{34}$  level
  - and ensure stable collisions of *Nanobeams* and preservation of their small emittance
- The second is useless if the first cannot be achieved, but is not less important



# LC Challenge 1: Energy



- Goal of 250 GeV/beam (and higher)
- Normal Conducting (JLC/NLC, CLIC) and
- Super Conducting (TESLA) RF technologies
  
- Teams are working hard to ensure successful jump from what is achieved, to the energy goal
  - SC technology - must jump from achieved 1 GeV (factor of 250)
  - NC technology - must jump from achieved 50 GeV (factor of 5)

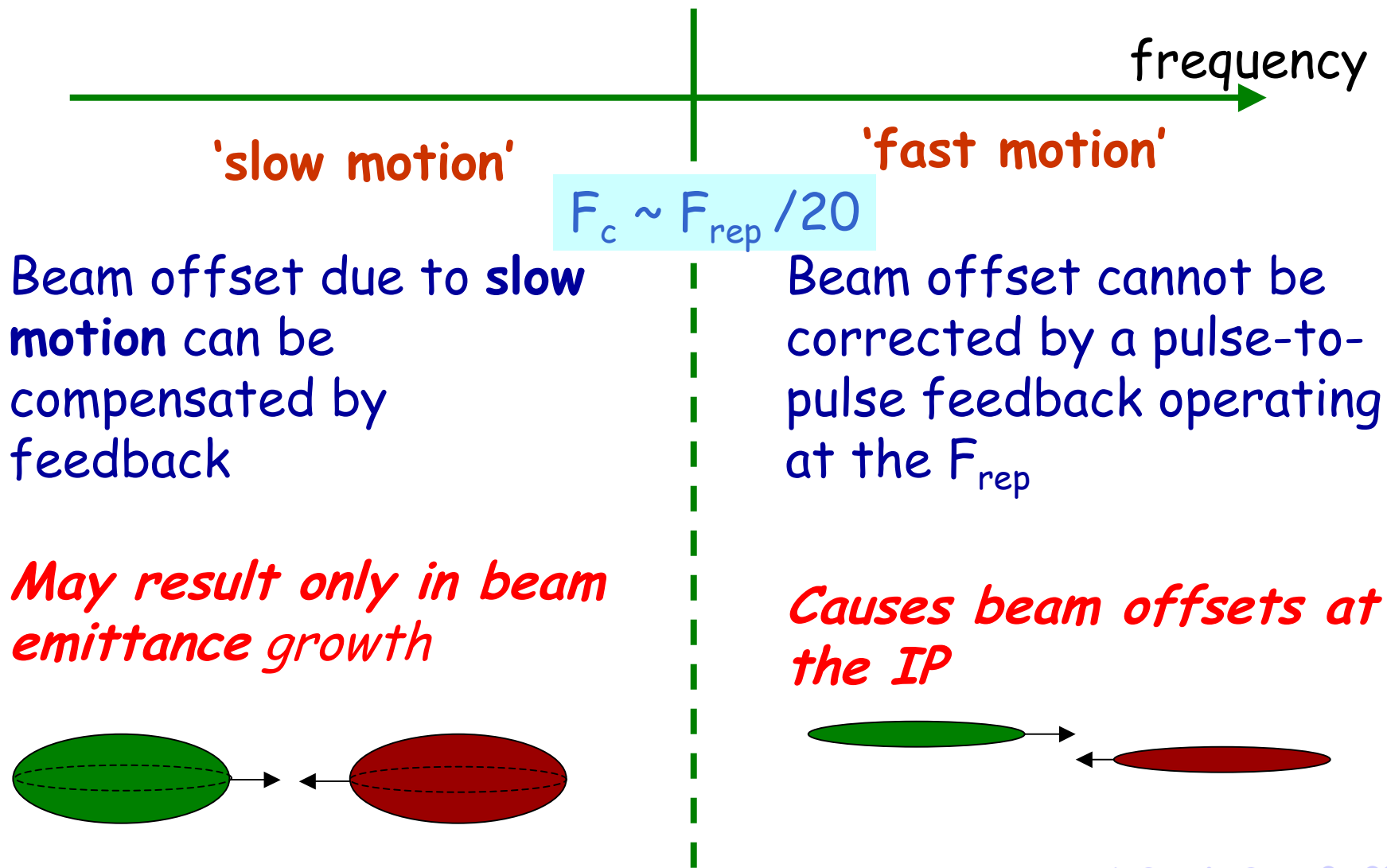
Significant progress along this way in the recent years



## LC Challenge 2: Luminosity

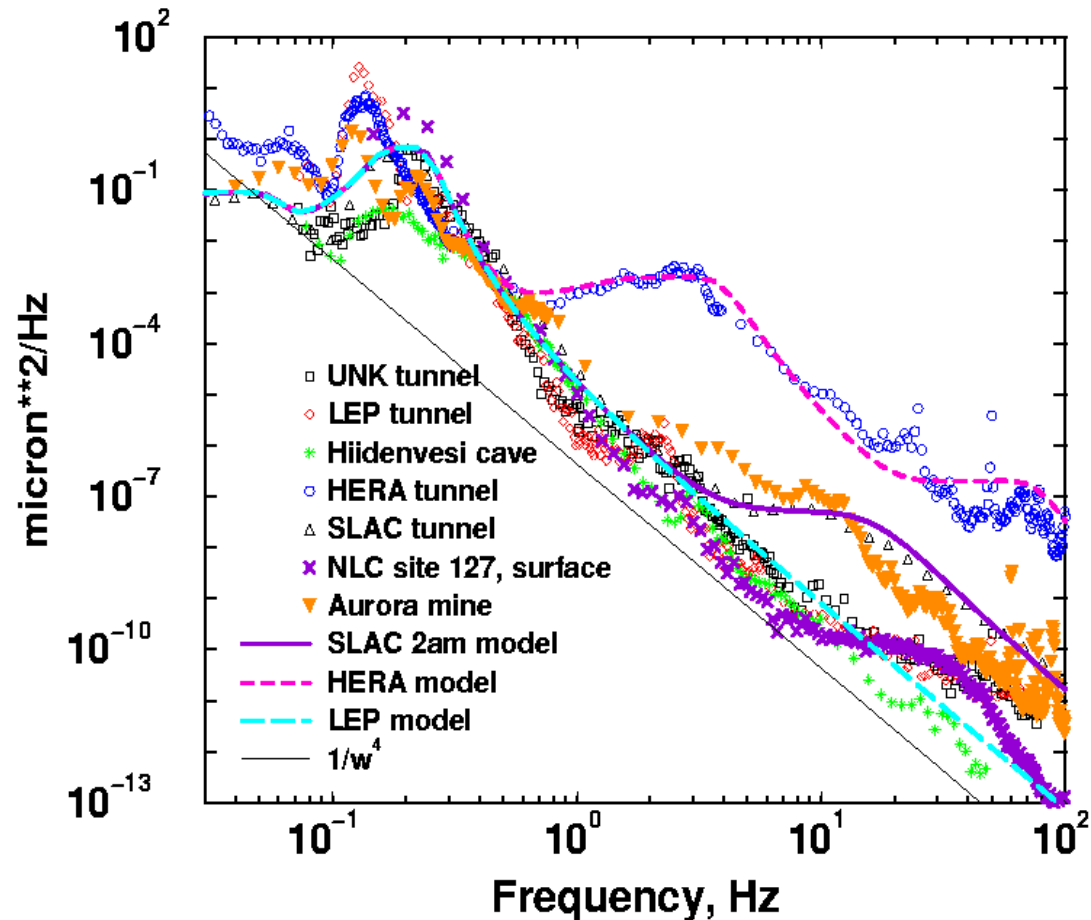
- Must jump by a Factor of 10000 in Luminosity !!!  
(from what is achieved in the only so far linear collider SLC)
- Many improvements, to ensure this : generation of smaller emittances, their better preservation, ...
- And need to provide stability
  - I.e. ensure that ground motion, remotely and locally created vibrations do not produce intolerable misalignments of LC elements

# Two effects of ground motion in Linear Colliders



# Evaluating effects of ground motion and vibration

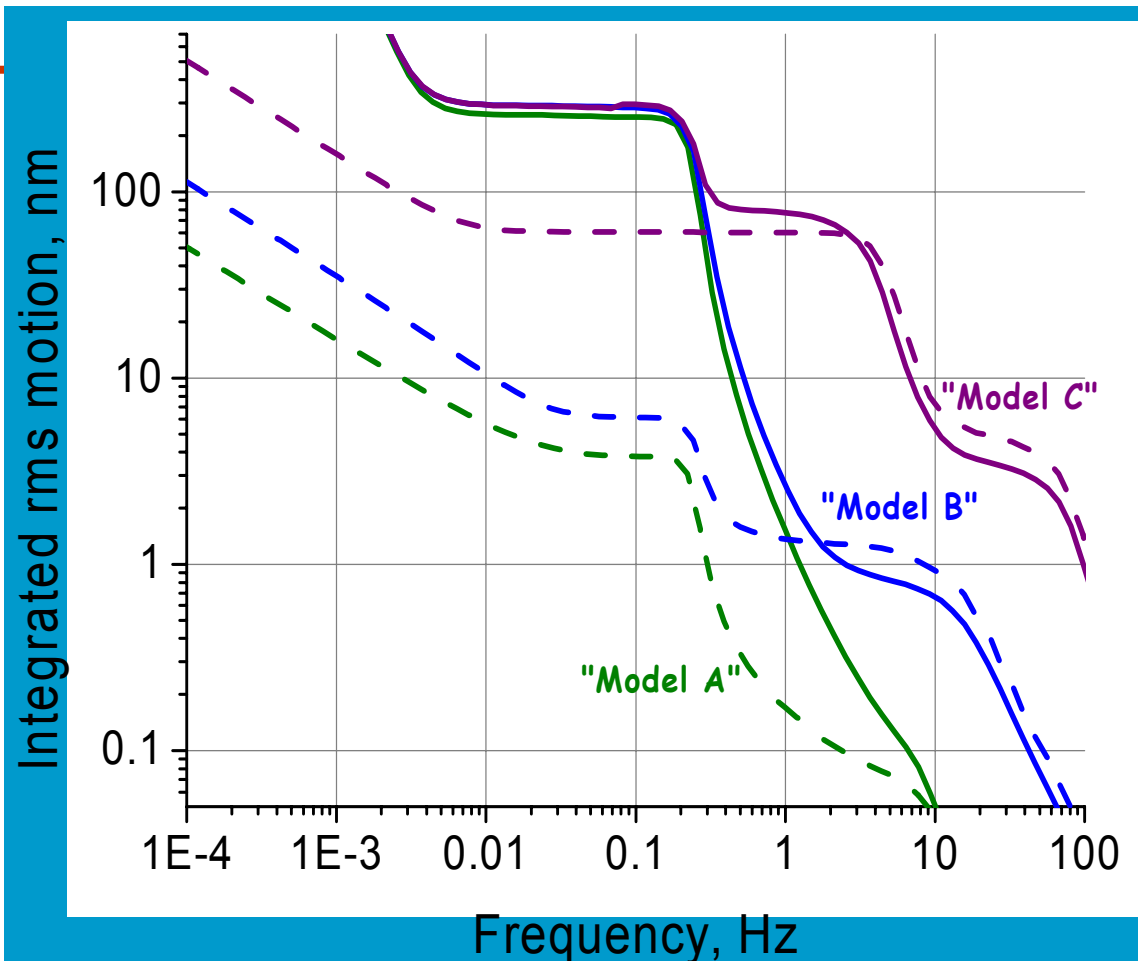
- Collect and understand data on ground motion and vibrations
- Build a model(s) of ground motion (e.g.  $P(\omega, k)$  spectrum)
- Then make simulation how LC performs
  - Apply corrections, feedbacks, optimize them...
- Decide whether this ground motion or parameters of LC are acceptable



Data from different locations  
1989 - 2001

# Ground motion models

- Based on data, build modeling  $P(\omega, k)$  spectrum of ground motion which includes:
  - Elastic waves
  - Slow ATL motion
  - Systematic motion
  - Cultural noises



Example of integrated spectra of absolute (solid lines) and relative motion for 50m separation obtained from the models

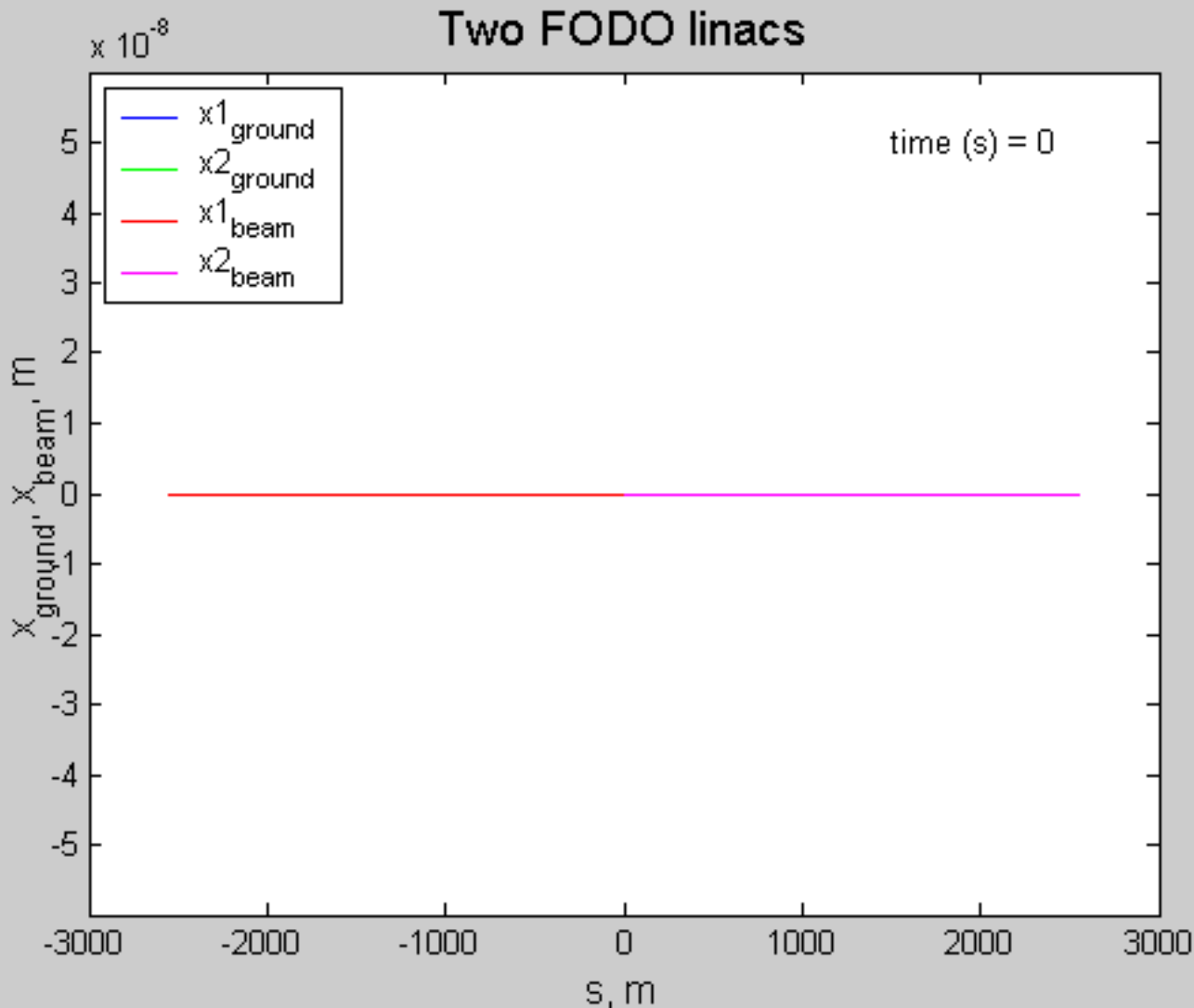




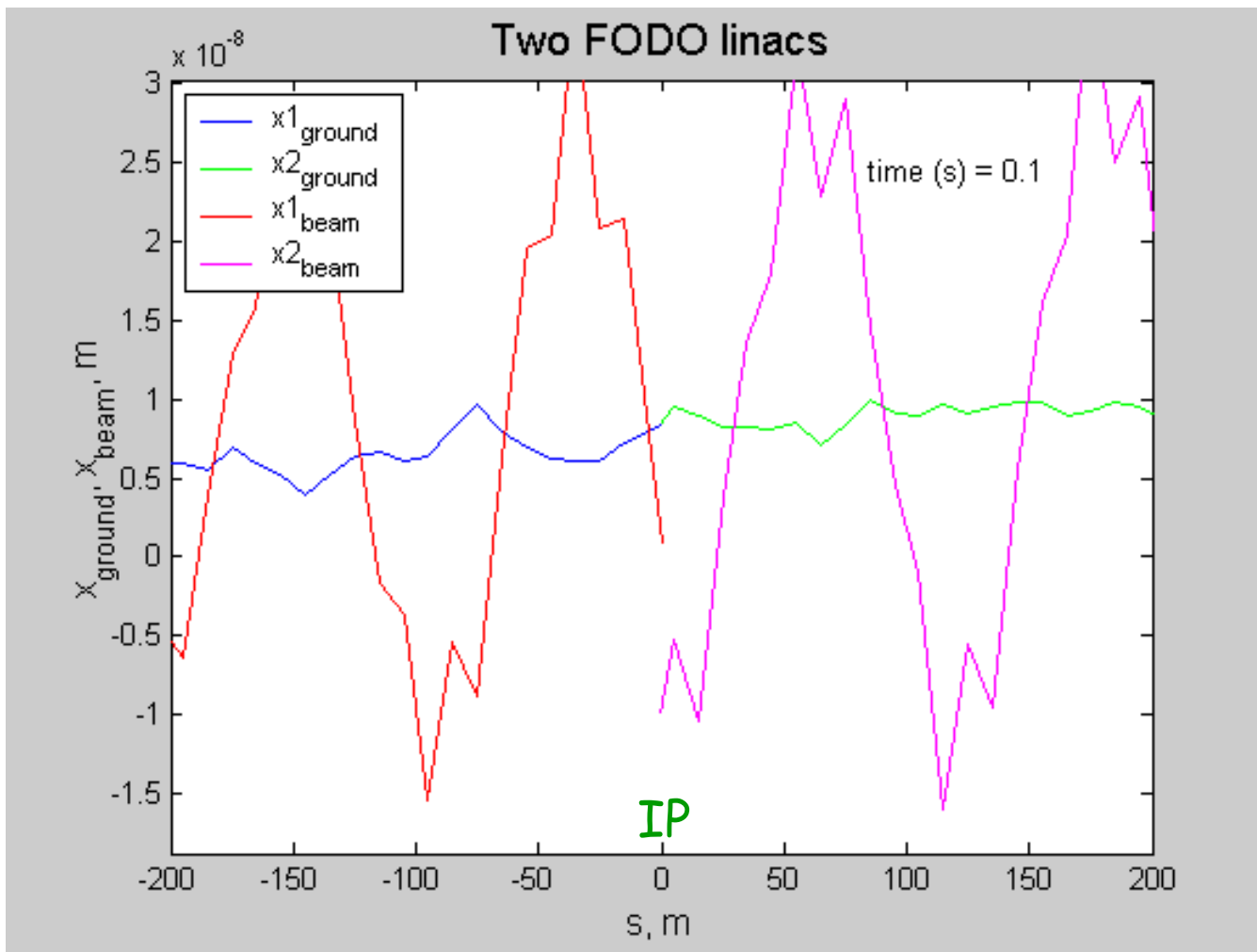
## Caution

- We should not forget that
  - Quads are not imbedded in a rock, but are sitting on supports or in cryostats
  - There are noise sources just on girders (e.g. from cooling water)
- Even if ground motion is acceptable, it is very important to verify, that stability of collider elements is sufficient
  - Further in the talk (and later during Workshop) we will discuss ongoing R&D that should answer this question

# Example: effect of ground motion on two FODO linacs pointing to each other



# Important that correlation between $e^+$ and $e^-$ beamlines is preserved



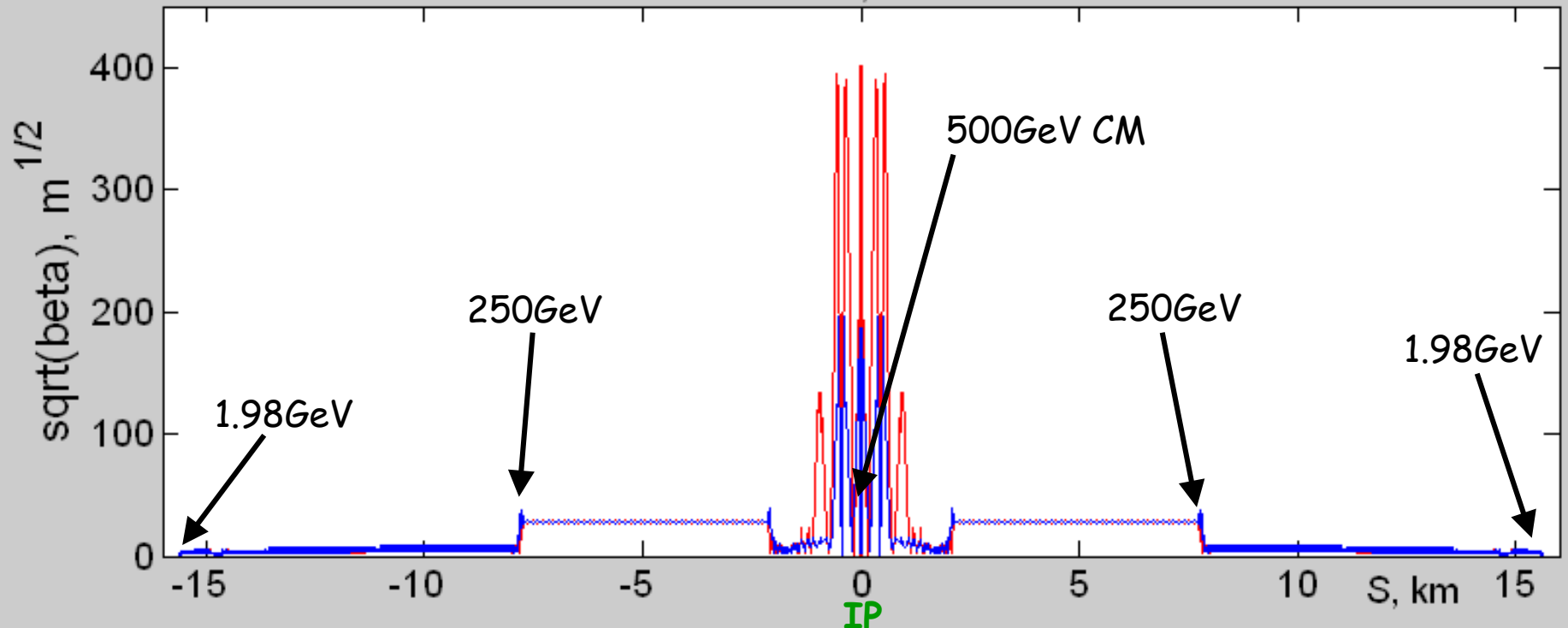
Note that ground is continuous, but beams have separation at the IP



# Simulations of complete NLC

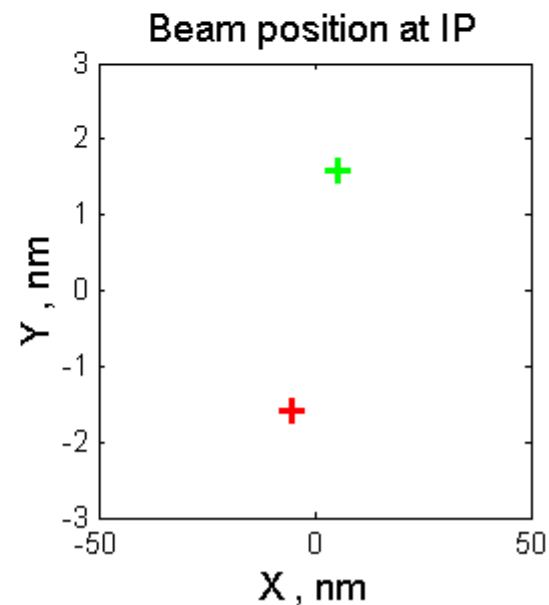
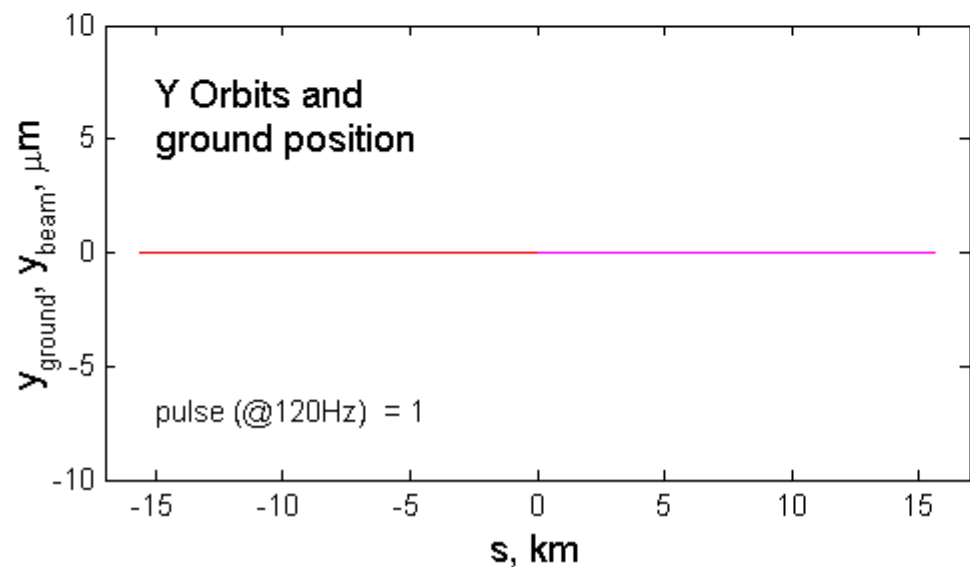
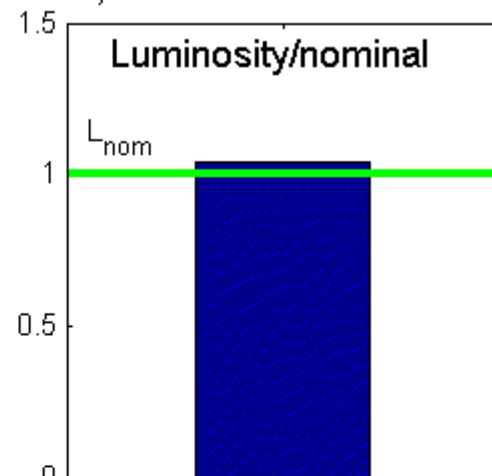
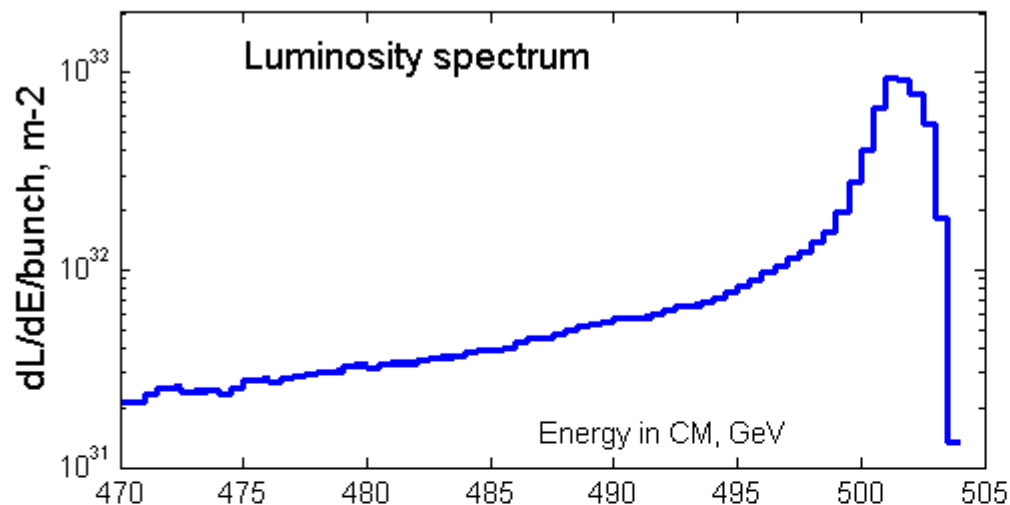
## DR => IP <= DR

NLC beta-functions, e+ & e- beamlines



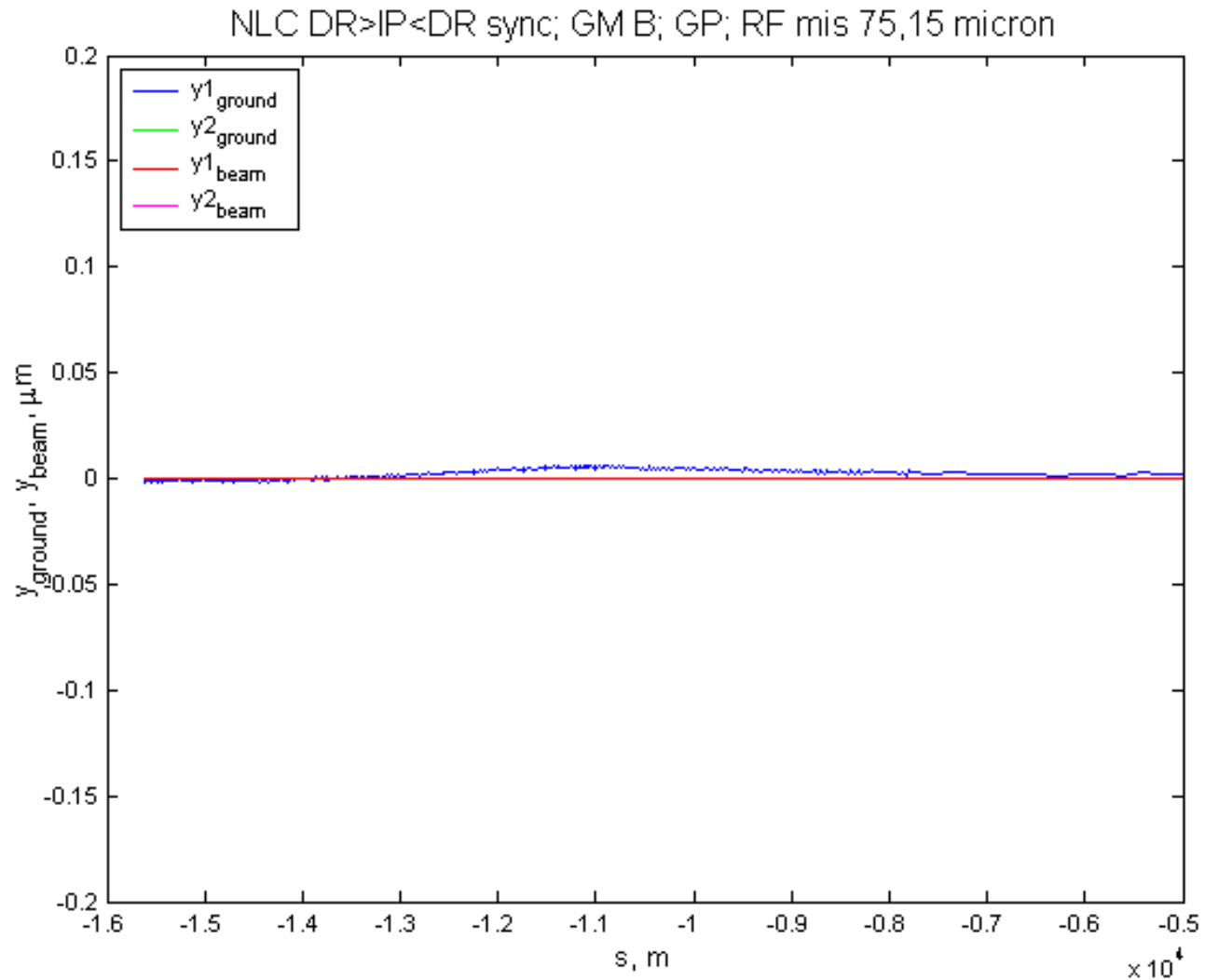
- Included:
- ground motion
  - train-to-train IP feedback
  - Errors in the linac
  - Beam-beam effects ...

NLC, DR>IP<DR; GM B; RF misal(x,y)=75,15 microns, IP feedback



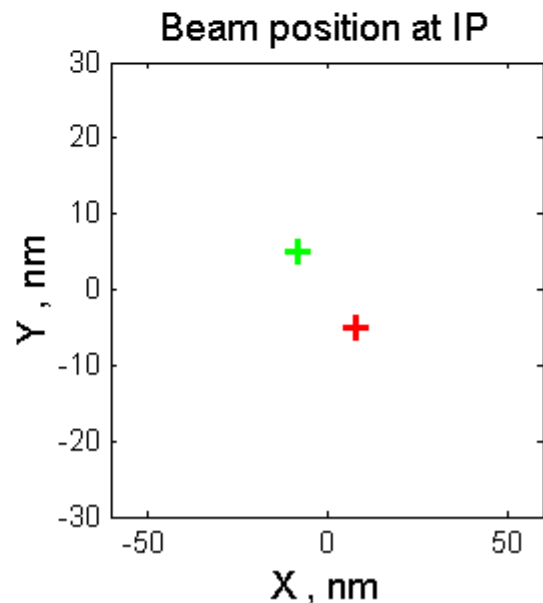
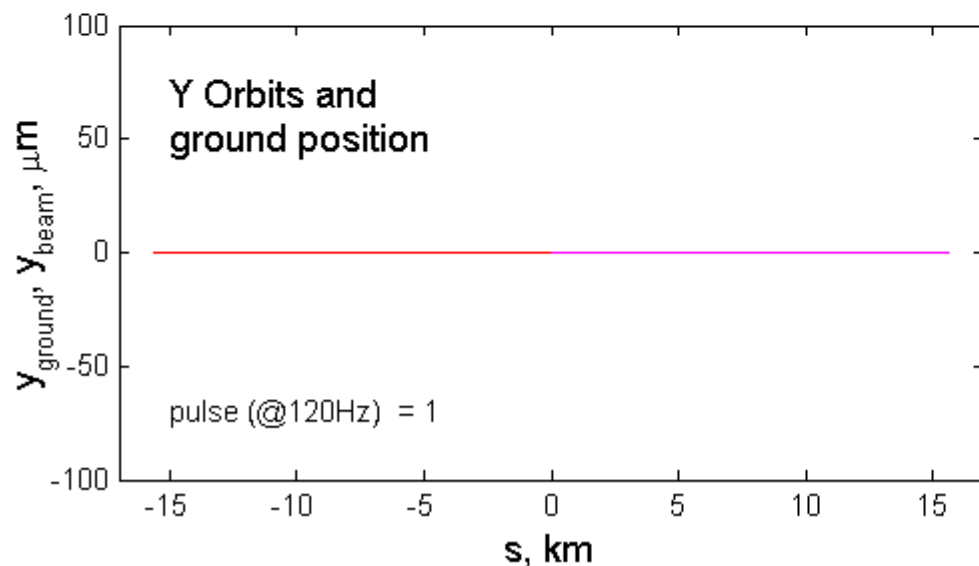
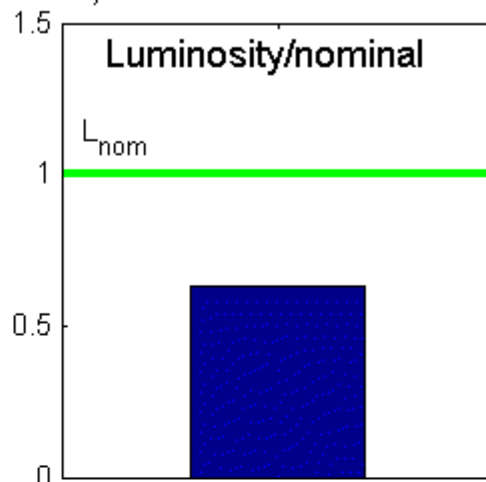
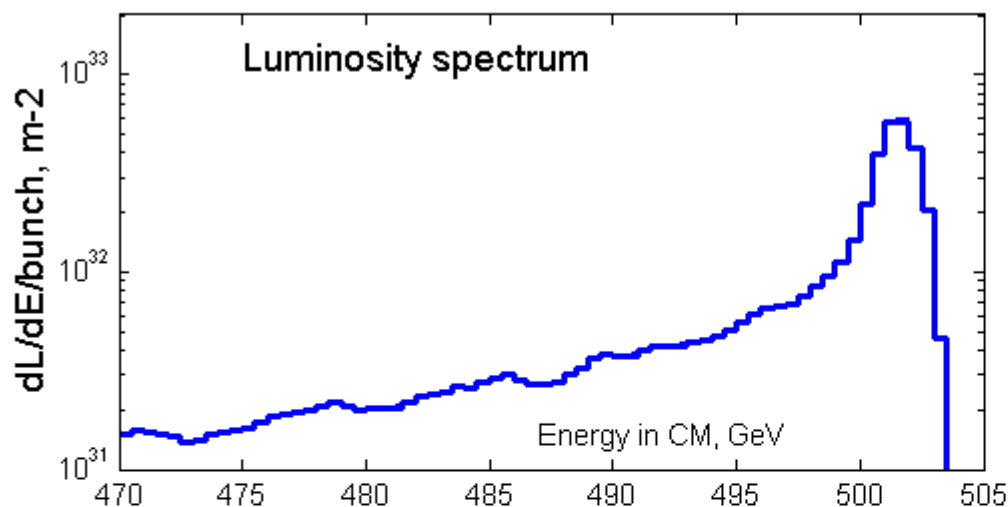


# Zoom into beginning of e- linac ...



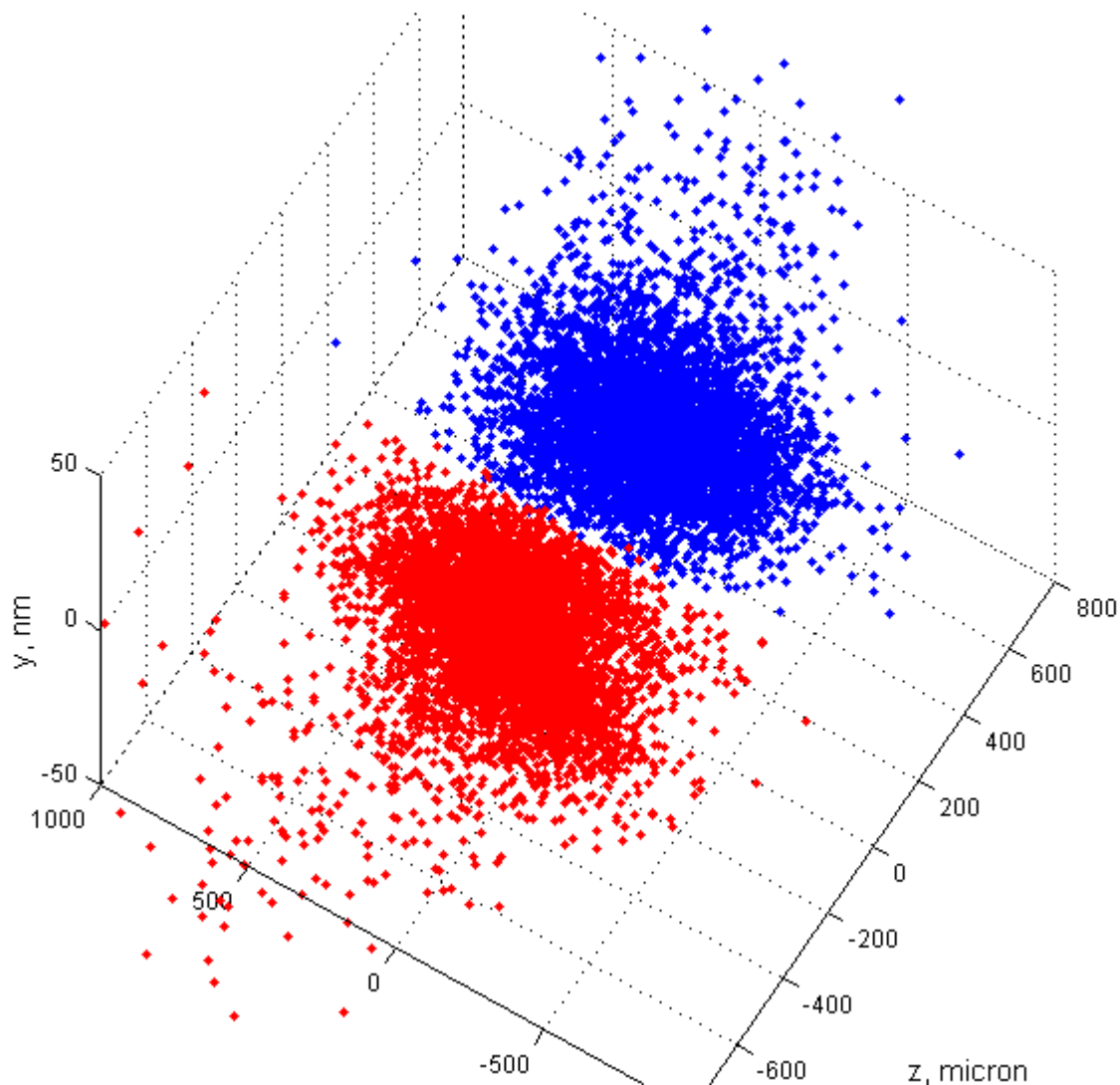
Transition from linac to transfer line

NLC, DR>IP<DR; GM C; RF misal(x,y)=75,15 microns, IP feedback

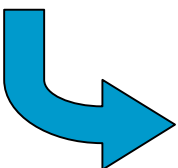




# Beam-beam collisions calculated by Guinea-Pig [Daniel Schulte]



"Banana effect" is included

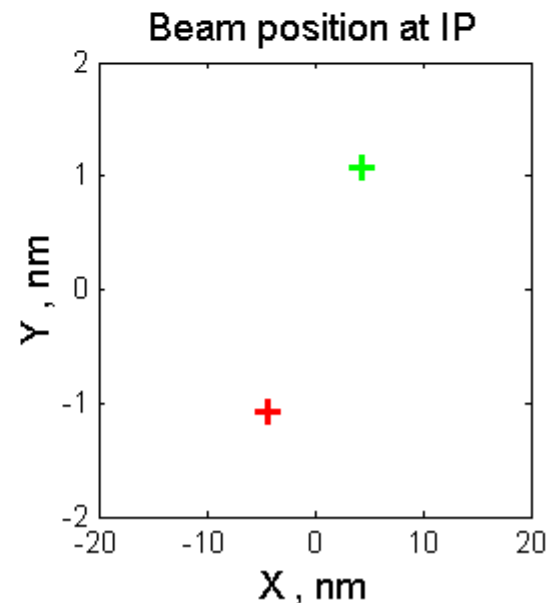
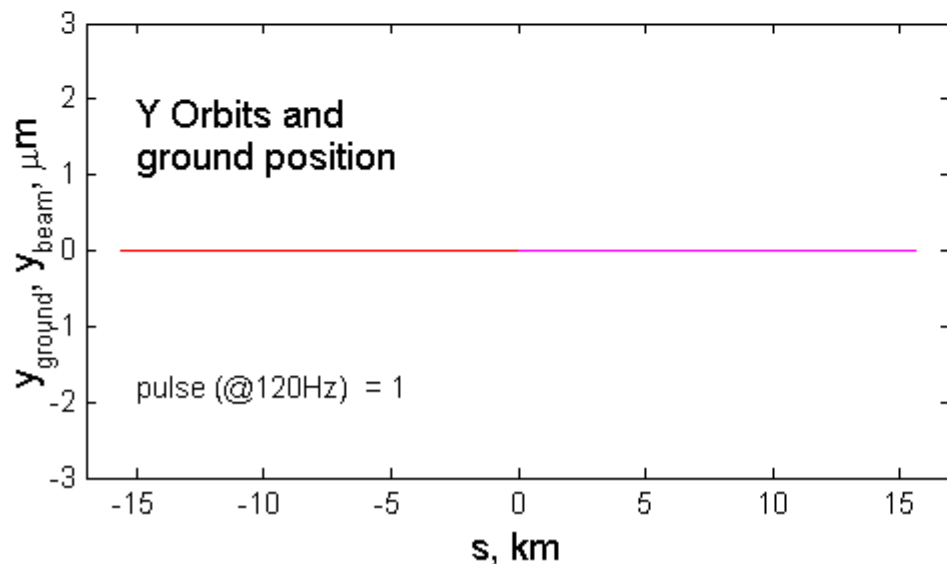
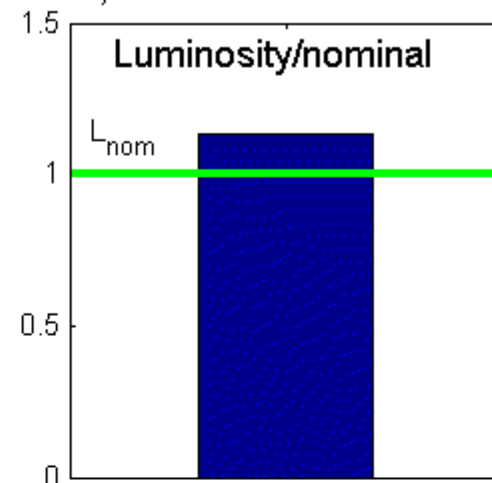
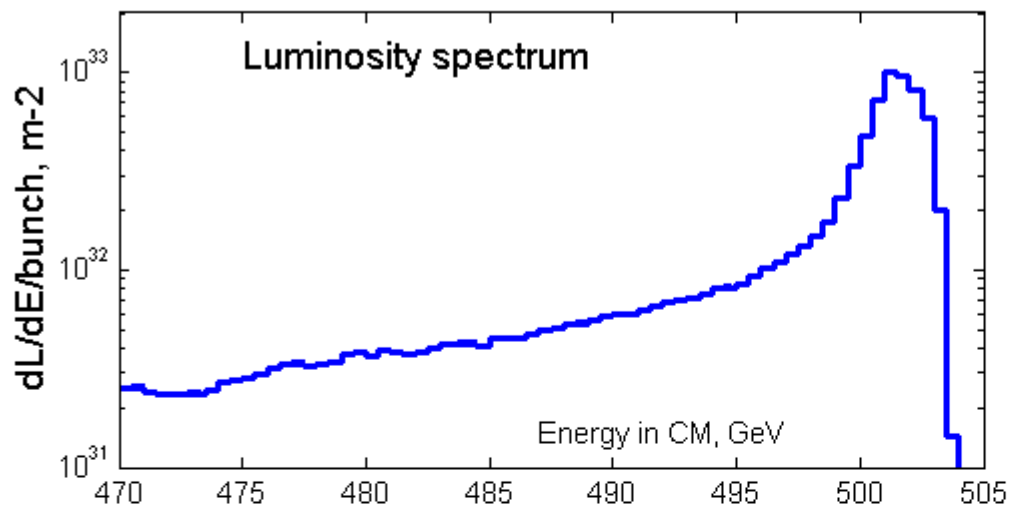


Daniel's talk





NLC, DR>IP<DR; GM A; RF misal(x,y)=75,15 microns, IP feedback





# IP beam-beam feedback

Colliding with offset  $e^+$  and  $e^-$  beams deflect each other

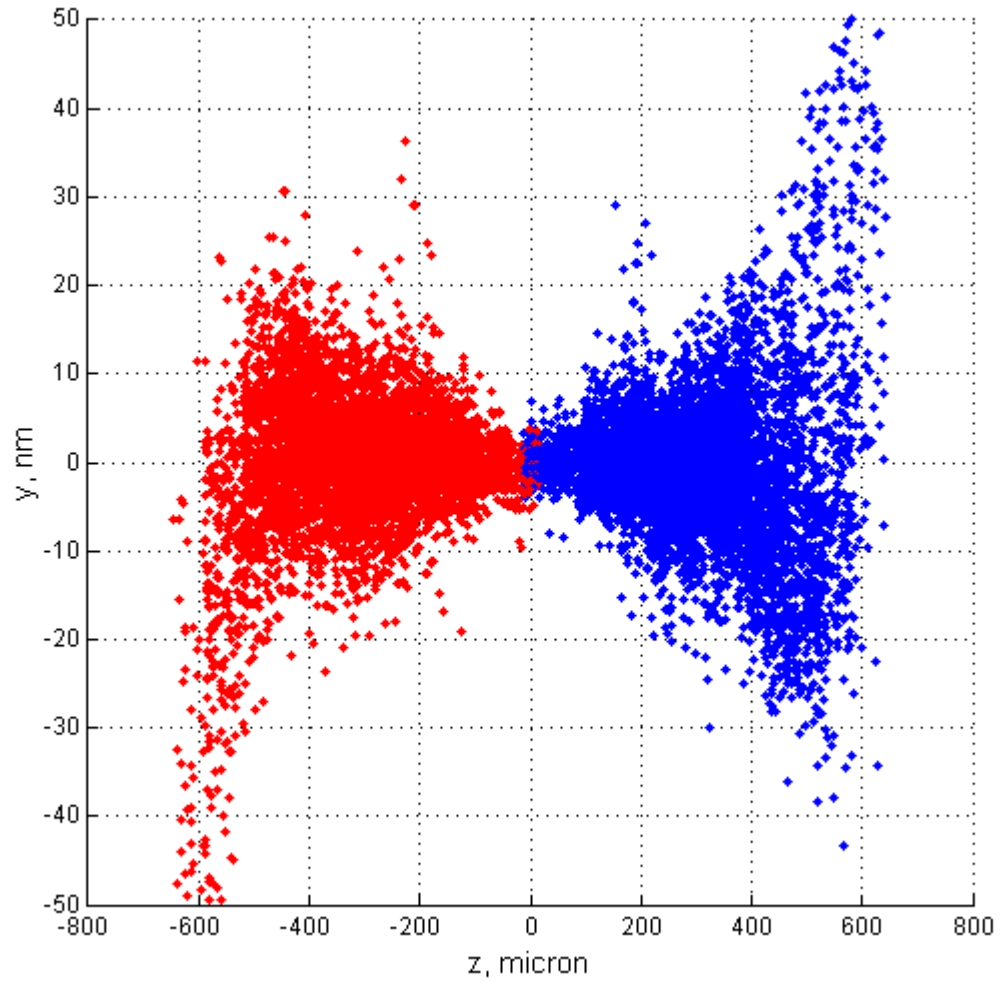
Deflection is measured by BPMs

Feedback correct next pulses to zero deflection  
(it uses state space, Kalman filters, etc. to do it optimally)

The previous page shows that feedback needs to keep nonzero offset to minimize deflection  
reason: asymmetry of incoming beams

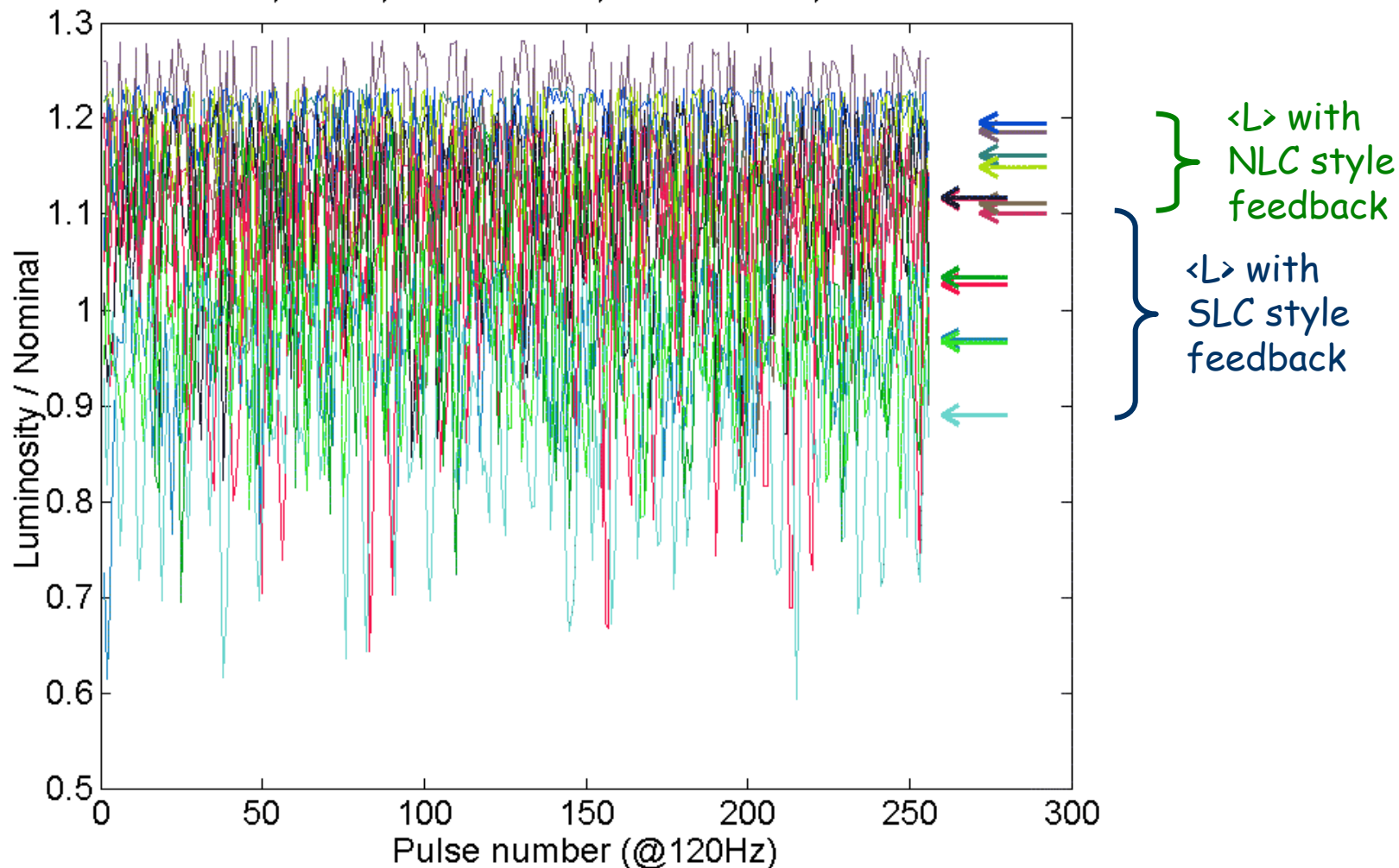
(RF structures misalignments  $\Rightarrow$  wakes  $\Rightarrow$  emittance growth)

# Pulse #100, Z-Y



# IP feedback developments and improvements

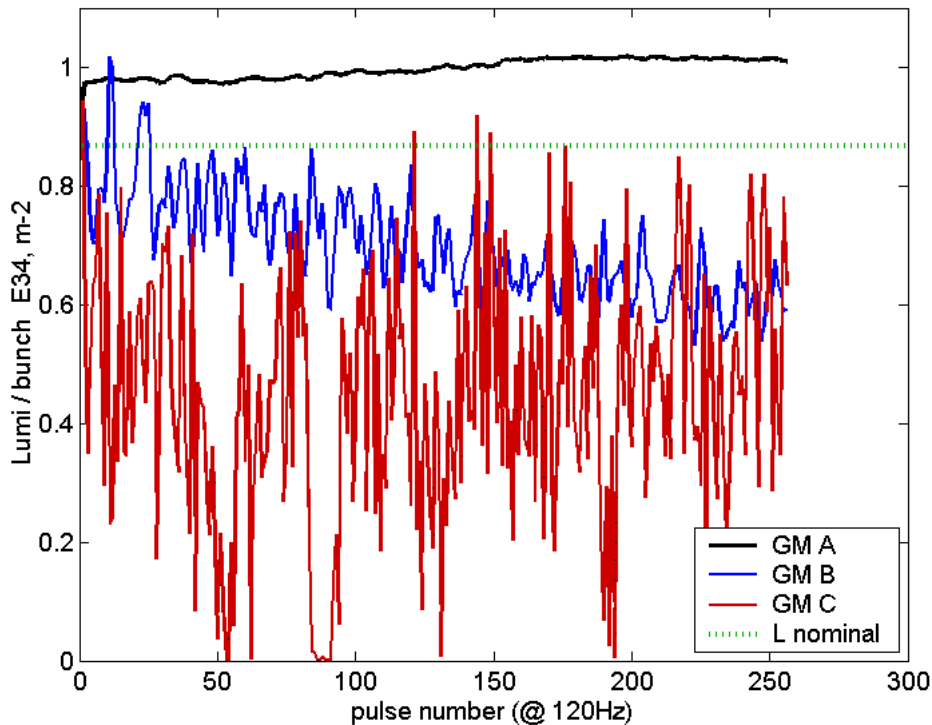
NLC, GM B, IP Feedback, RF misal=75, 15micron



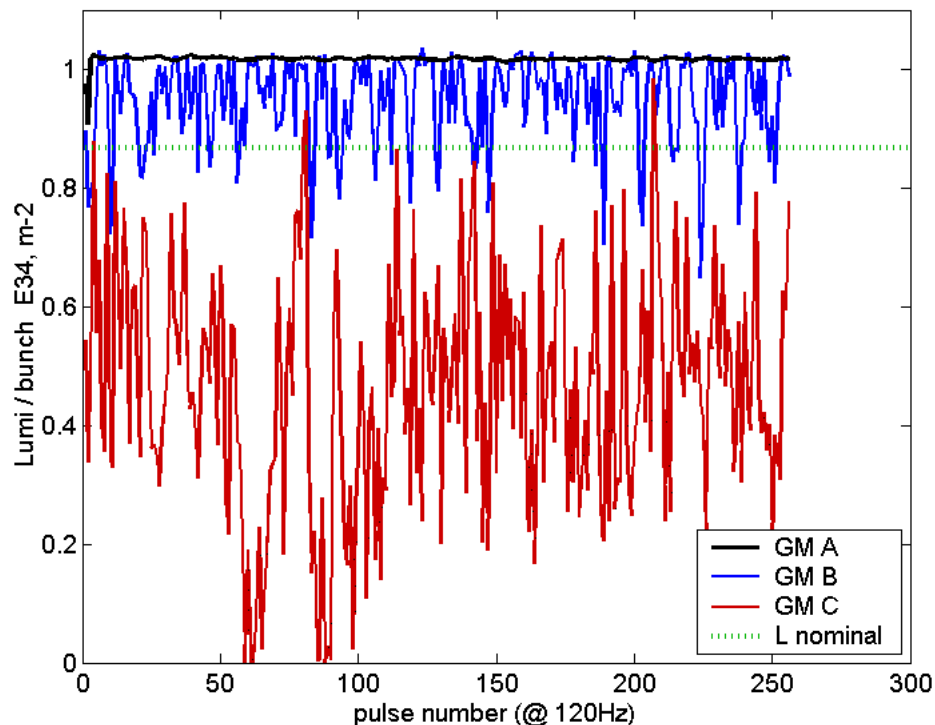
Talk of  
Linda Hendrickson

# With and without IP feedback, examples

NLC DR&gt;IP&lt;DR ; RFmisal(x,y)=75,15micron



NLC DR&gt;IP&lt;DR ; IP fdbk; RFmisal(x,y)=75,15micron



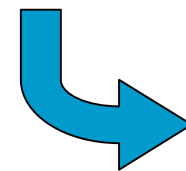
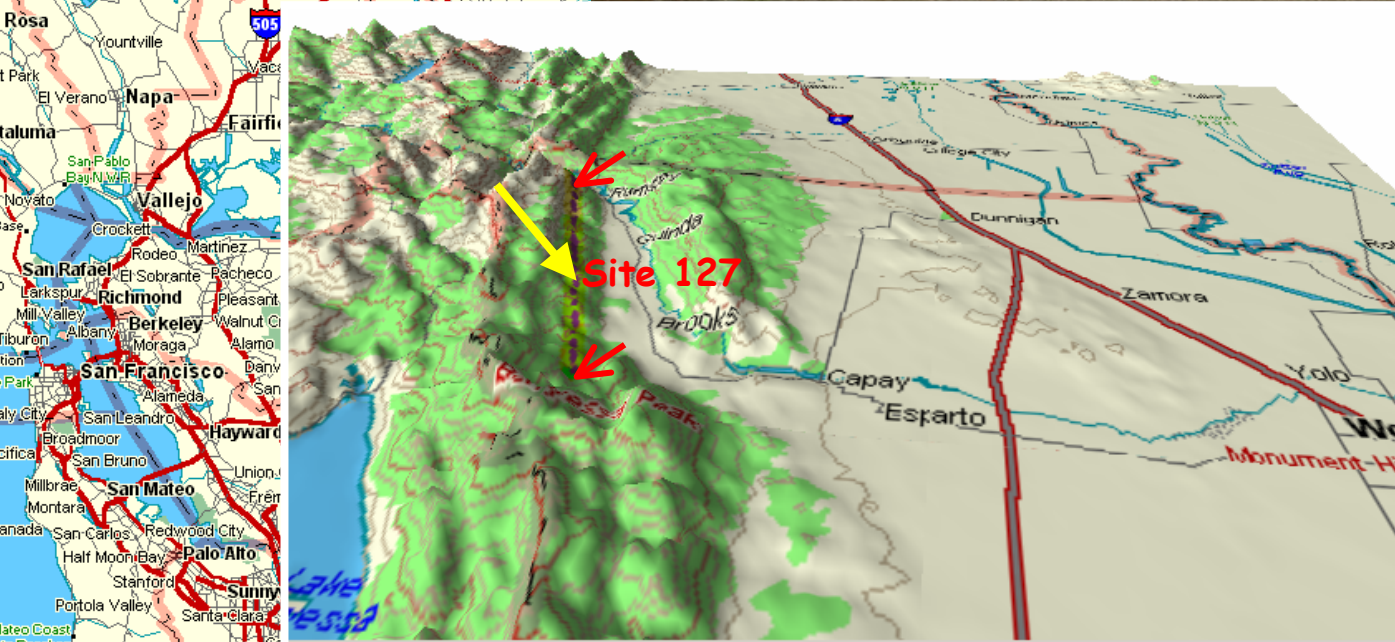
Example for one particular seed  
(seed is the same for the left and right plots)



# Ongoing and required R&D

- Studies of the sites stability
- Studies of near-tunnel noises and vibration transfer from the surface
- Studies of in tunnel noises, including vibration transfer from the parallel tunnel
- Studies of on-girder (in-cryostat) noises

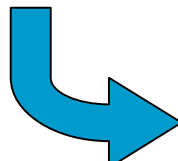
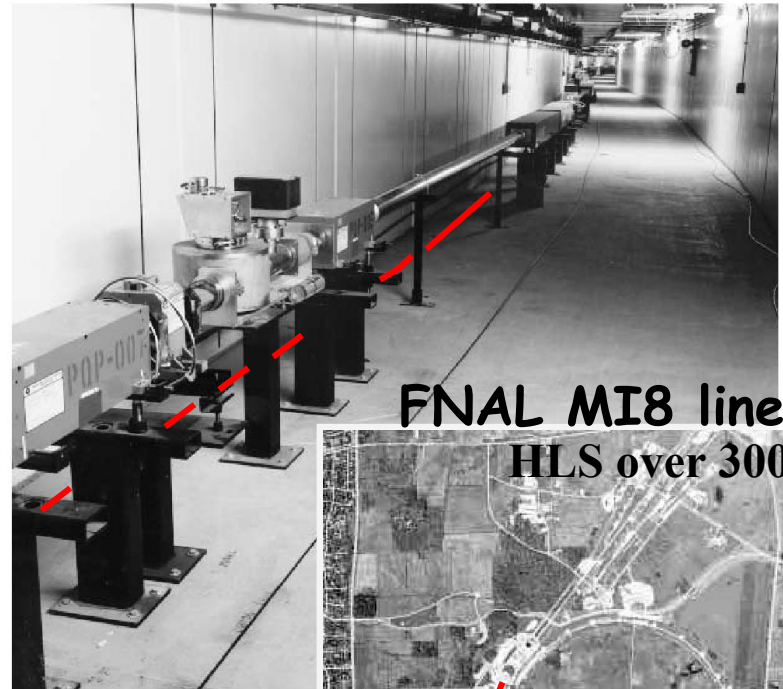
# Stable NLC sites in CA



Talk of  
Fred Asiri

A.Seryi, Sept.2, 2002

# BINP-FNAL-SLAC slow motion studies and HLS R&D

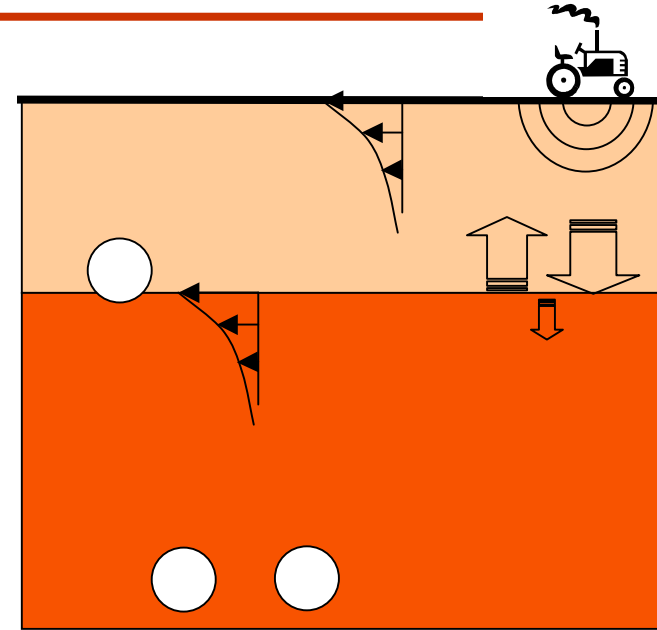
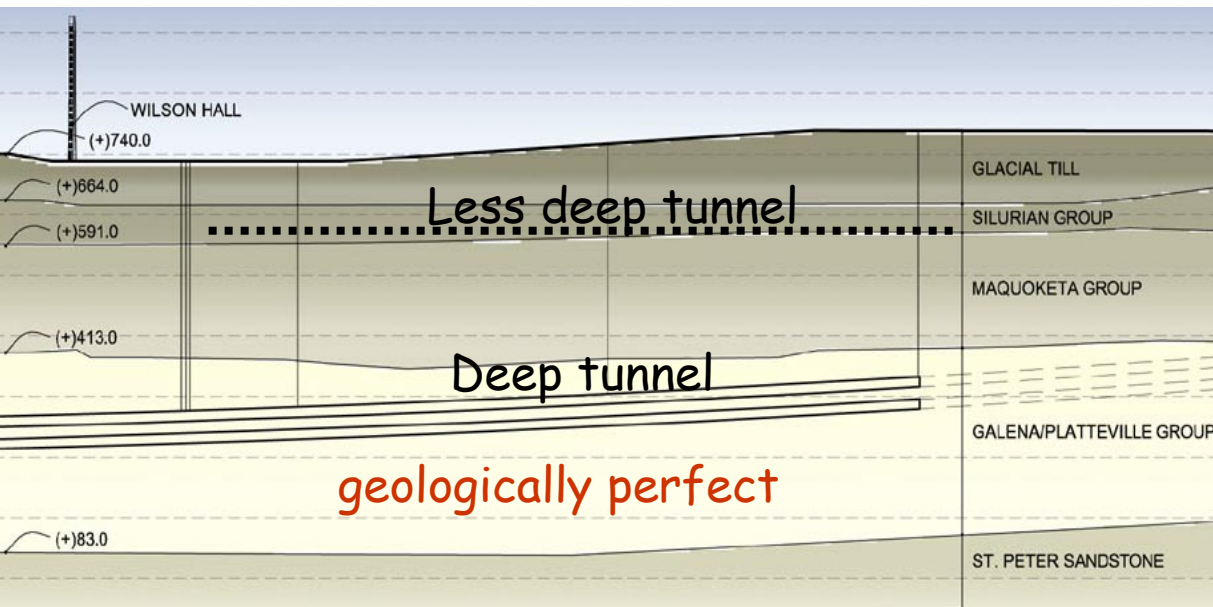


Talk of  
Vladimir Shiltsev

A.Seryi, Sept.2, 2002



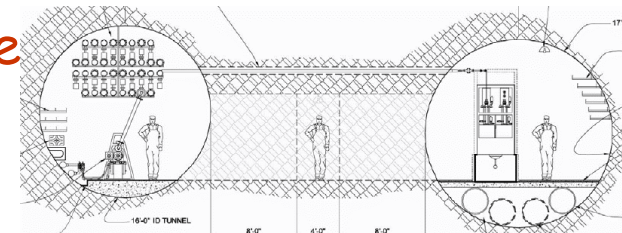
# Study of noise vs depth. Study of vibration transfer.



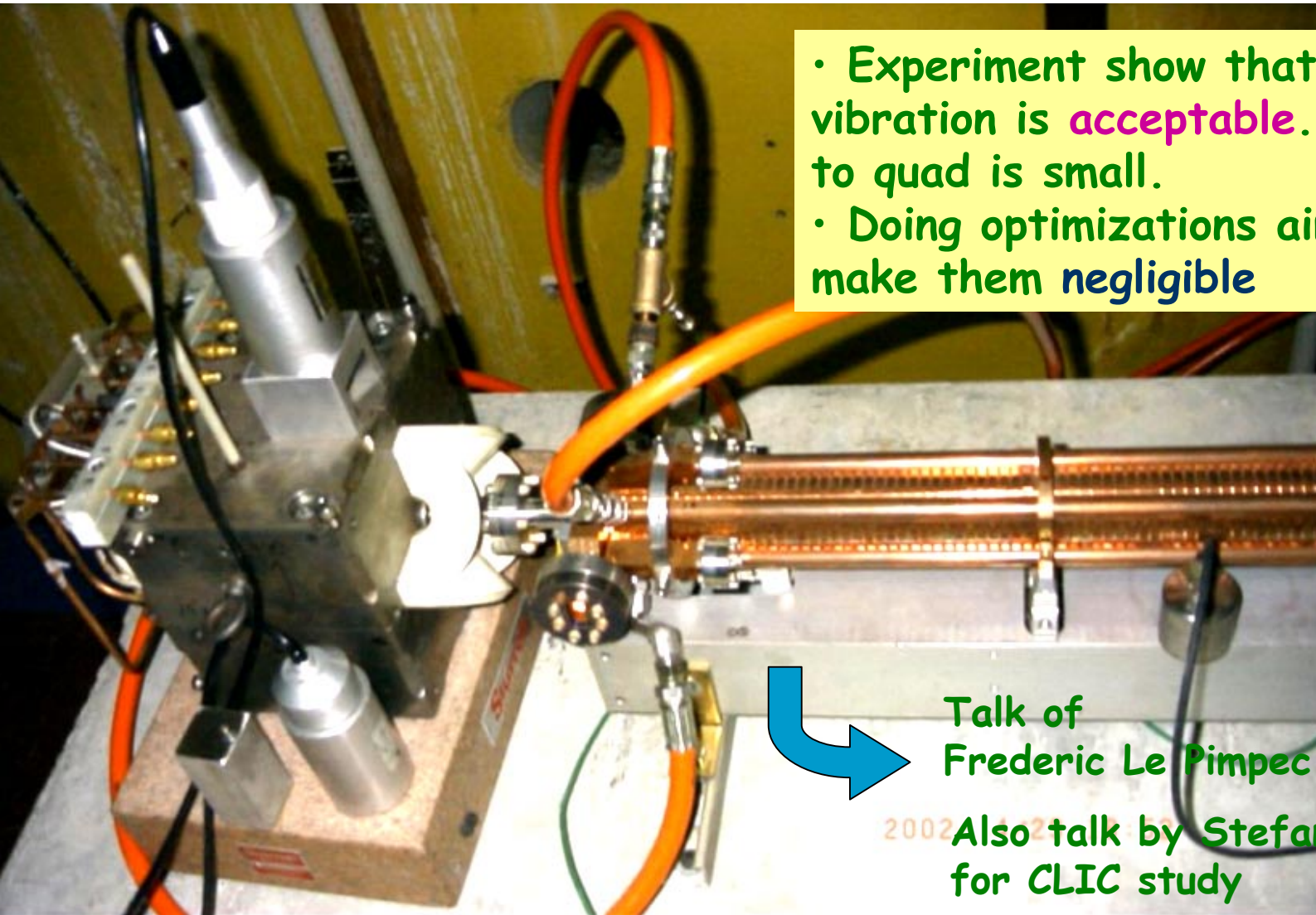
- Measurements in NUMI tunnel, noise vs depth dependence (FNAL and Northwestern Univ.)
- Vibration transfer from surface to shallow tunnel
- Plan to study vibration transfer between two parallel deep tunnels



Talk of  
Fred Asiri



# Vibration of RF structure due to cooling and vibration coupling to quadrupoles



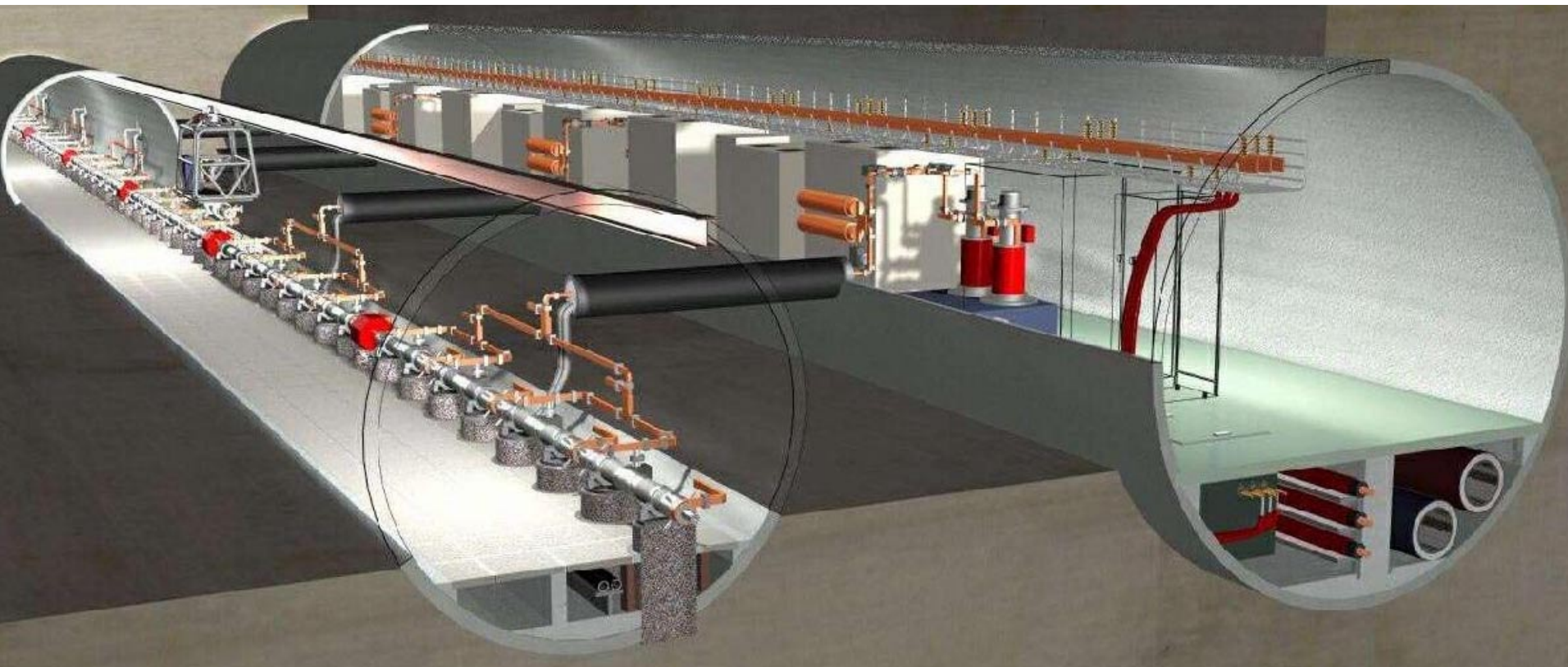
- Experiment show that additional vibration is **acceptable**. Coupling to quad is small.
- Doing optimizations aimed to make them **negligible**

Talk of  
Frederic Le Pimpec

2002-11-22 15:55  
Also talk by Stefano Redaelli  
for CLIC study

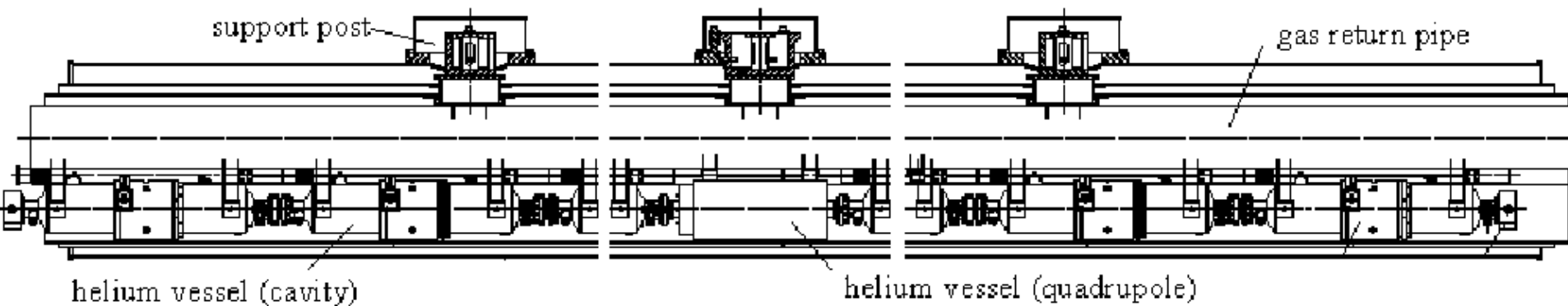
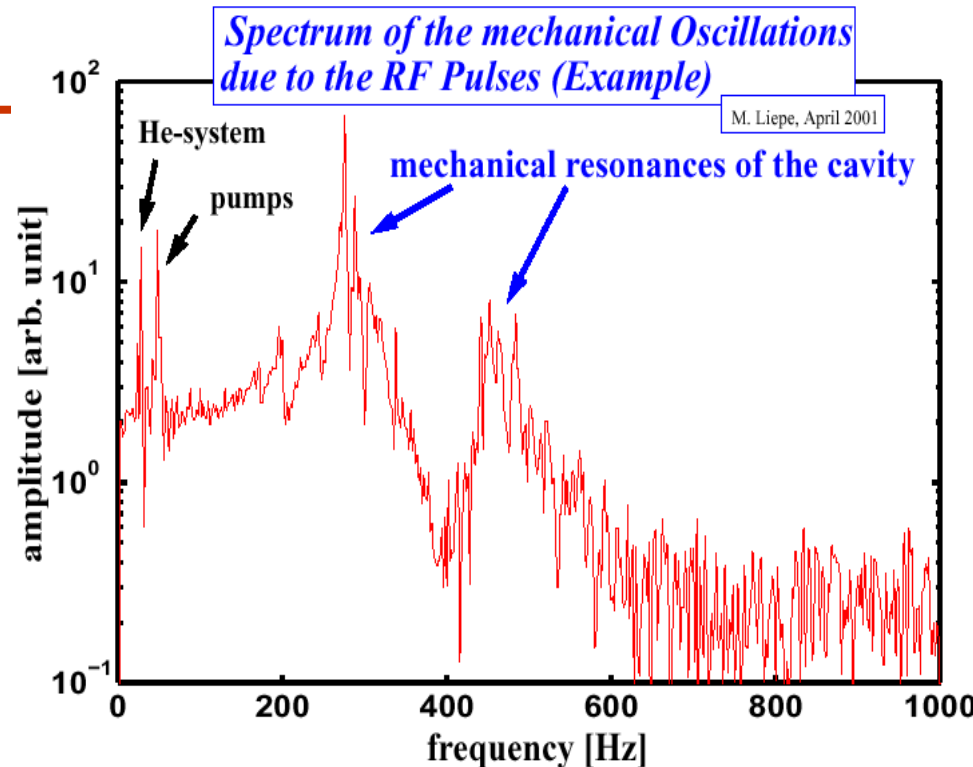
# Important feature of warm LCs: quads can have separate supports

- Quads on separate supports are connected to rock
- Vibration coupling from RF structure to quad can be made very small
- This helps to achieve vibration stability requirement for linac quads



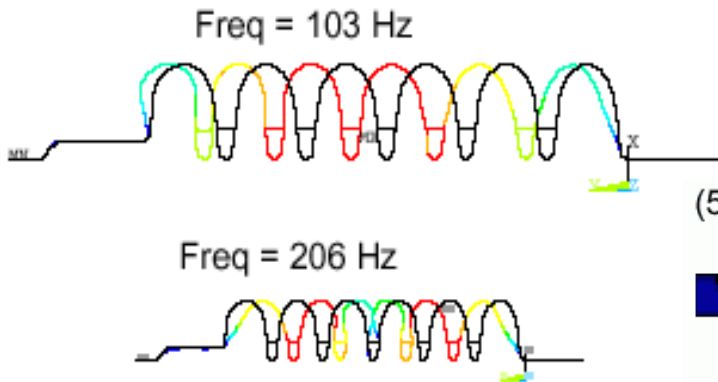
# Quad stability in TESLA linac

- Vibration stability requirement for SC linac are much looser than in warm LC
- Issue: common support (helium return pipe), which may be "a shaky ground"
- Noises: from RF pulse (Lorenz force); mechanical coupling to pumps, etc.
- Vibration coupling to quads need to be appropriately minimized by the design

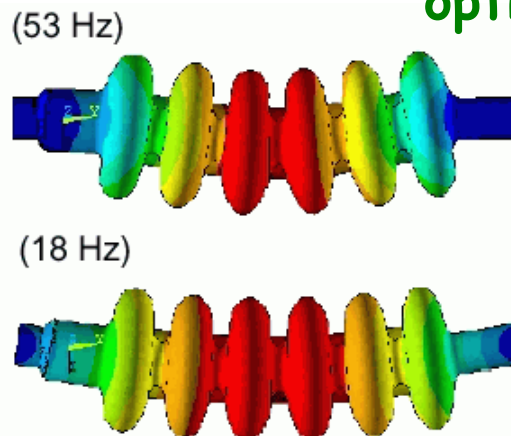


# Optimization of quad stability in SC linac

- There are a lot of experience with analysis and successful optimization of vibration properties of RF structures
  - To make it stiffer, optimize positions of supports, etc., so that to decrease detuning by RF pulse
- Similar techniques could be extended to optimize design to minimize quad vibration



**Example: Vibration modes of different SC cavities (for SNS) and their optimization [Carlo Pagani, Danilo Barni, SCPL 2000]**



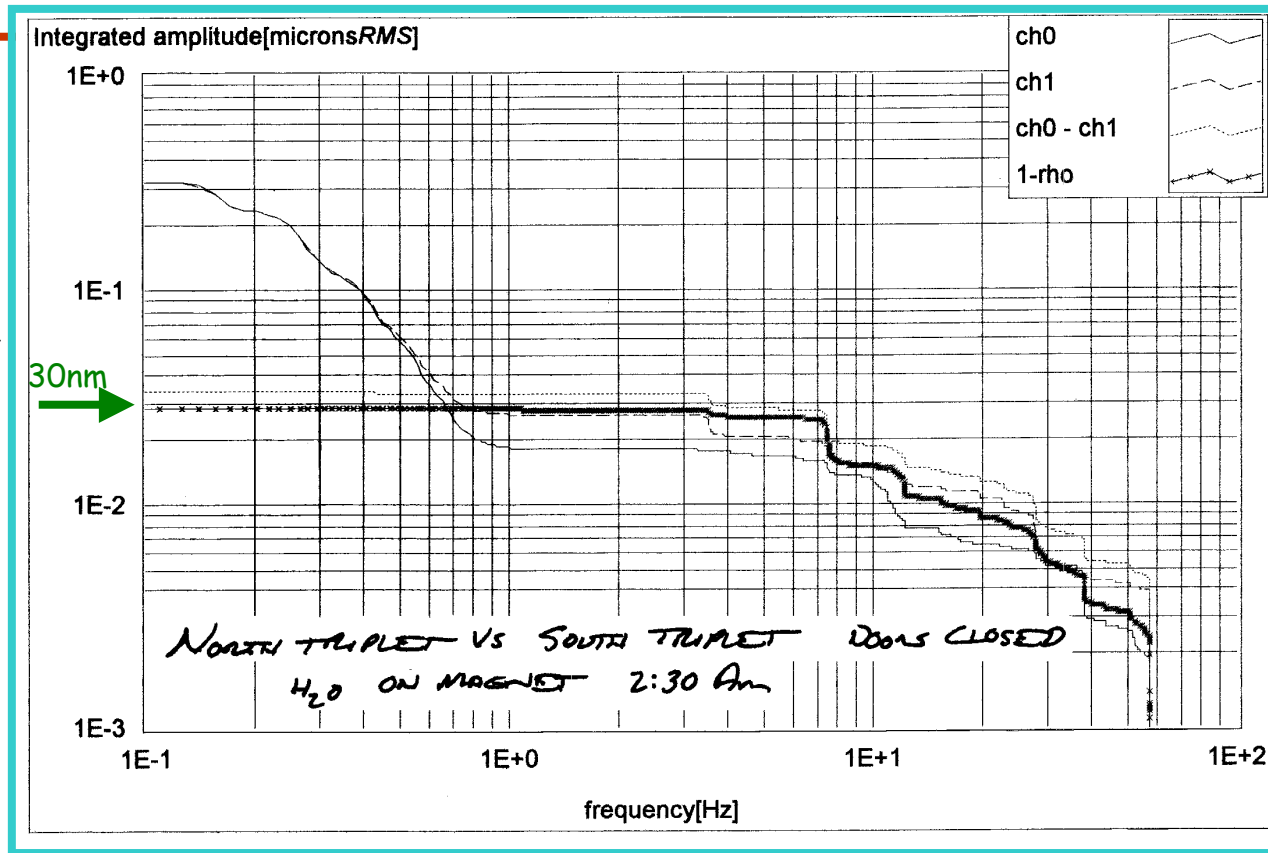
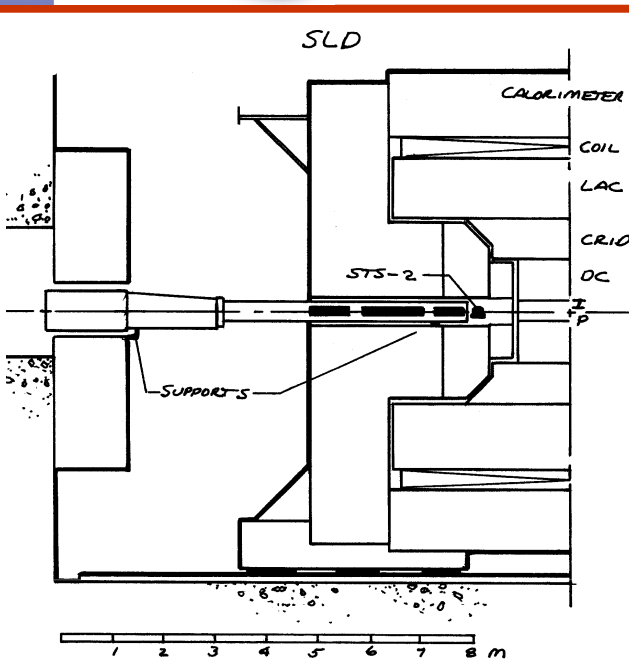


## Moving to the IP...

- Let's assume that we understand stability in linac
- And let's move our attention to the IP.  
What are stability problems there?
- FD has most stringent tolerances. And it may sit on a detector, which is "noisy ground"

# Cultural noise at detector

## 1995 SLD measurements [Gordon Bowden]

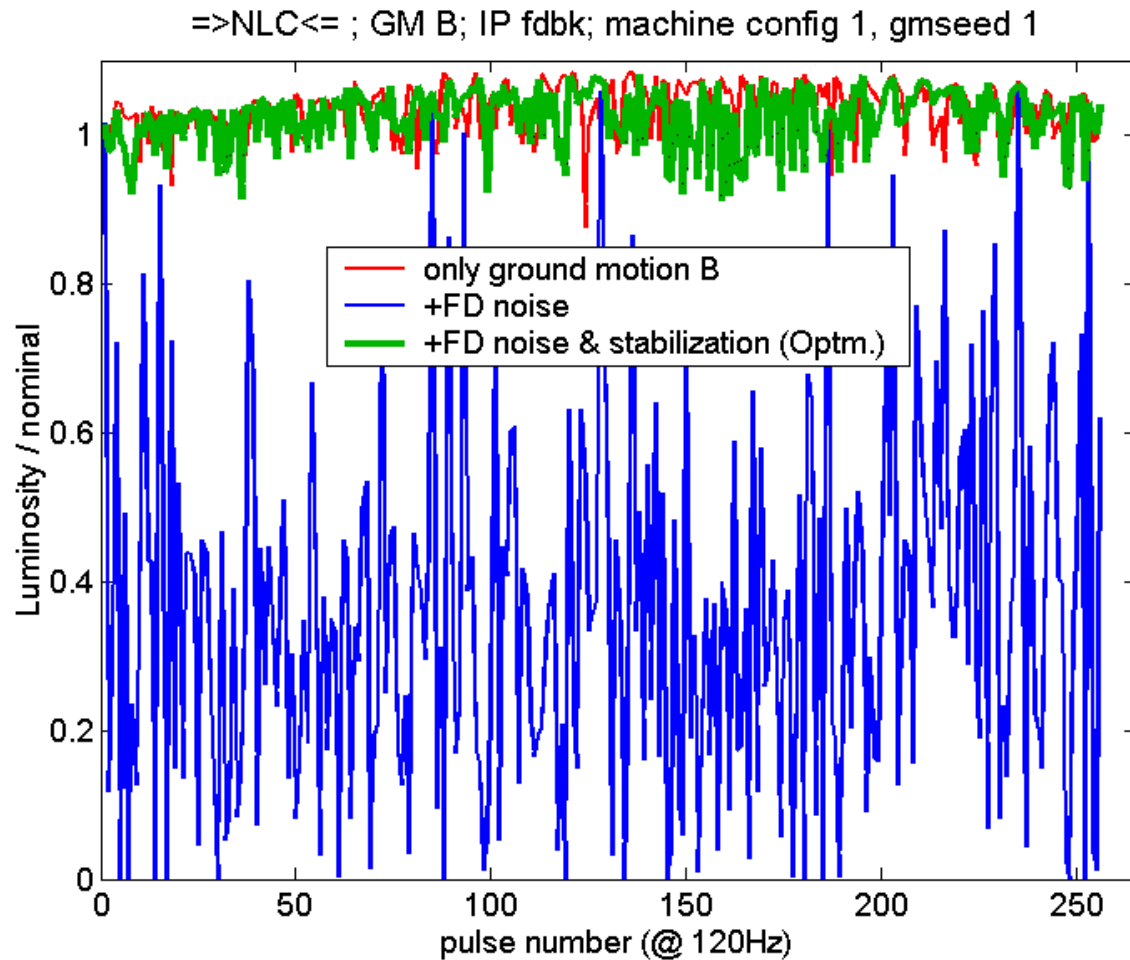


- Measured  $\sim 30\text{nm}$  relative motion between South and North final triplets  
Magnetic field was OFF (magnetic field ON could have increases detector rigidity). North triplet (Ch1) noisier - this side of the building is closer to ventilation and compressor stations. Resonances (3.5Hz, 7Hz) are likely to be resonances of detector structure.
- More quiet detector certainly possible.



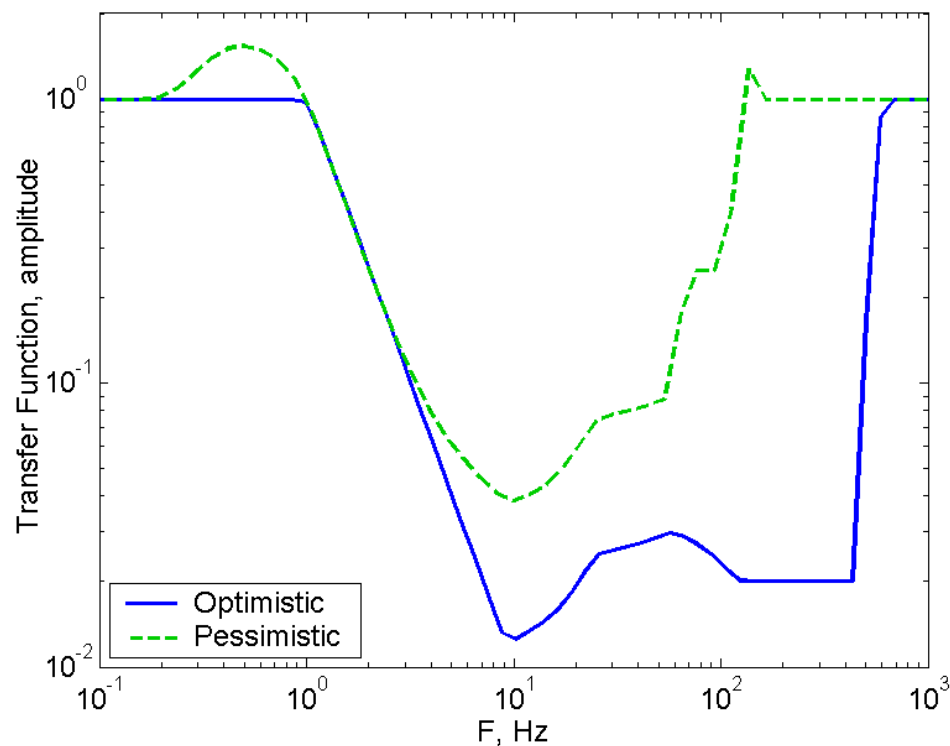
# Performance with and without FD stabilization

- Assume pessimistic, SLD-like FD vibration
- Then luminosity drops significantly (to  $\sim 1/3$ )
- If FD is actively stabilized or corrected, luminosity is restored

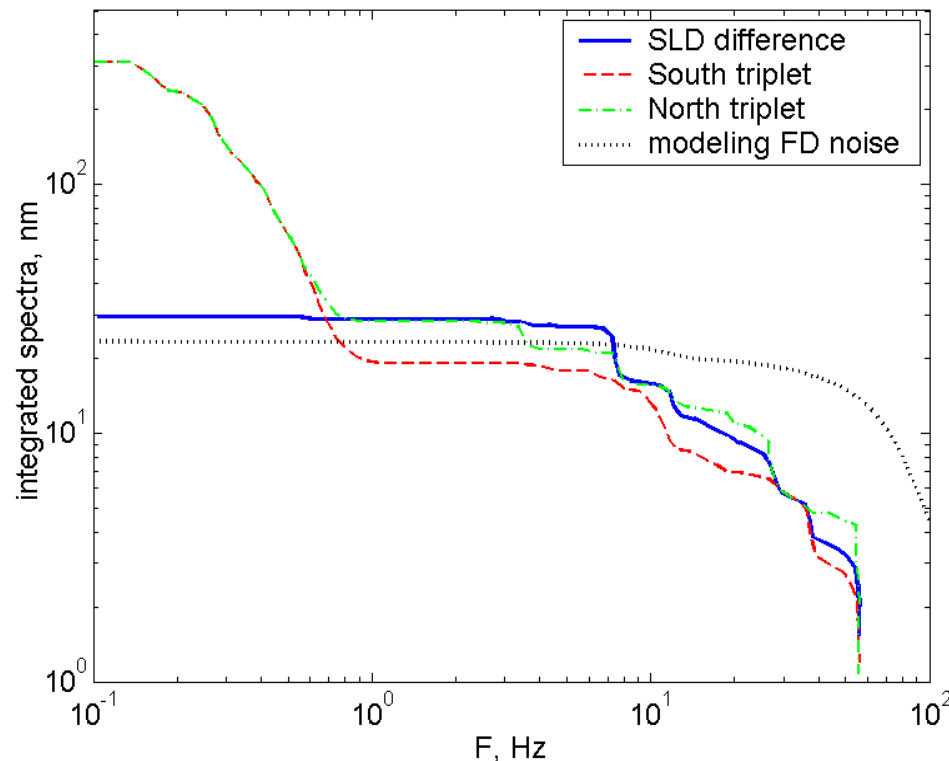




# FD stabilization modeling assumption



FD active stabilization (correction) is represented by Transfer Functions. Optimistic and pessimistic curves. The curves do not necessarily imply a particular stabilization or correction choice.

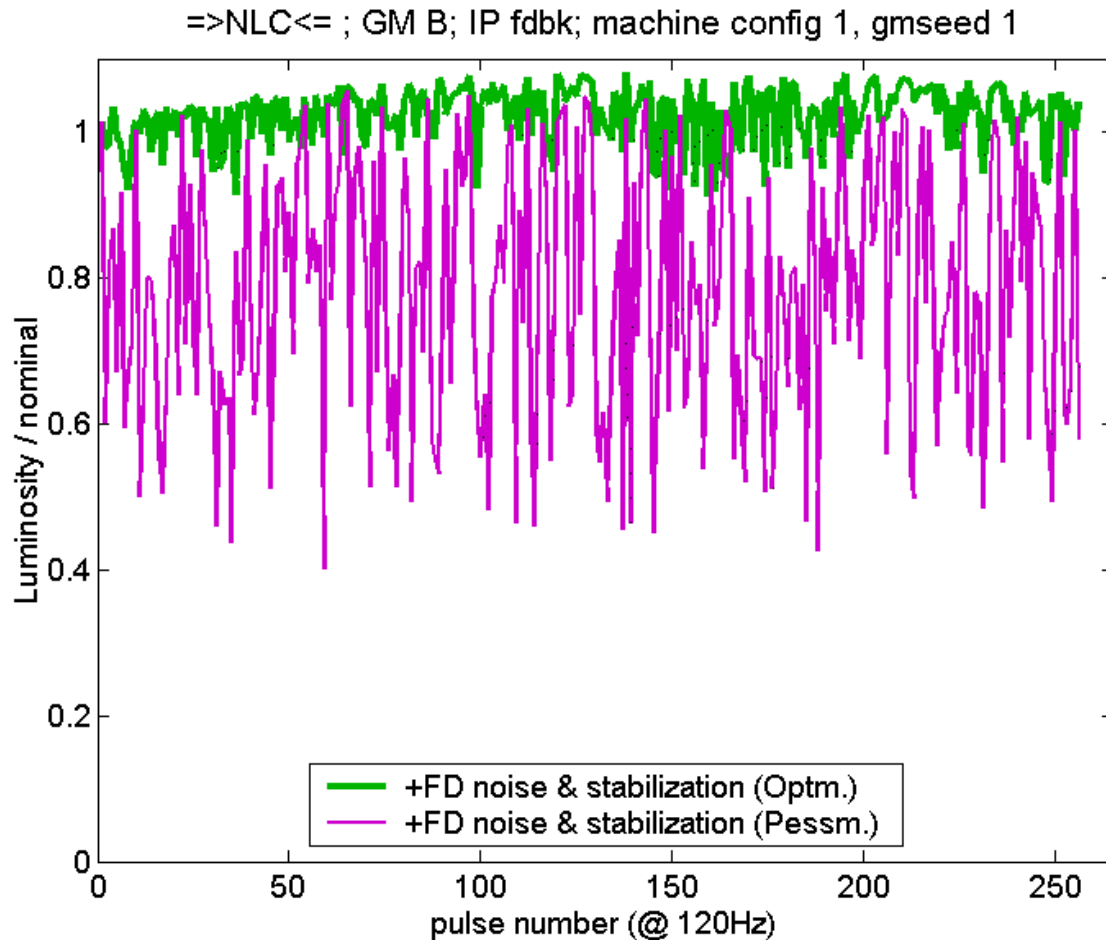


Noise measured at SLD [Bowden,95] and FD noise modeling spectrum. Same amplitude as in SLD is pessimistically assumed. The noise is shifted to higher frequencies (assuming the detector structural resonances are improved).

# Performance with different optimism about FD stabilization

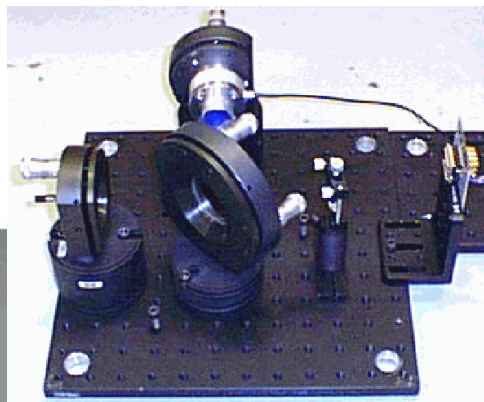
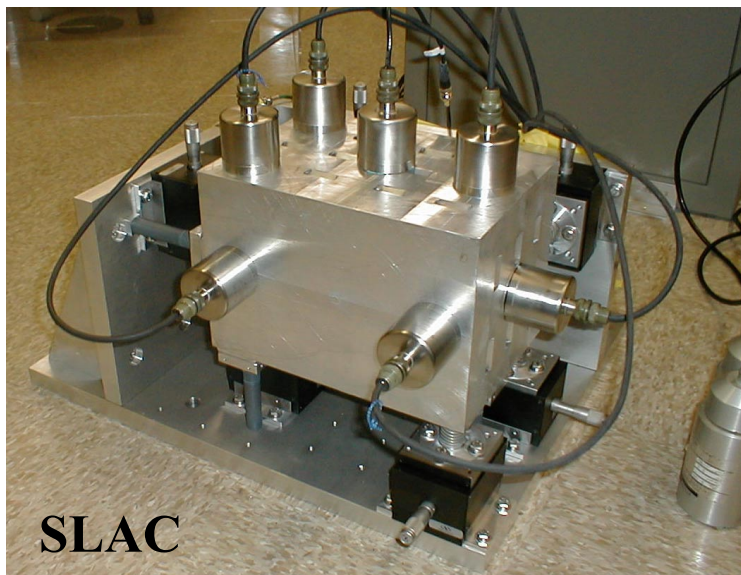
- With optimistic FD stabilization (correction) performance the luminosity is restored almost completely (<1% reduction)

- With pessimistic stabilization (correction) performance, the reduction of luminosity is ~25%



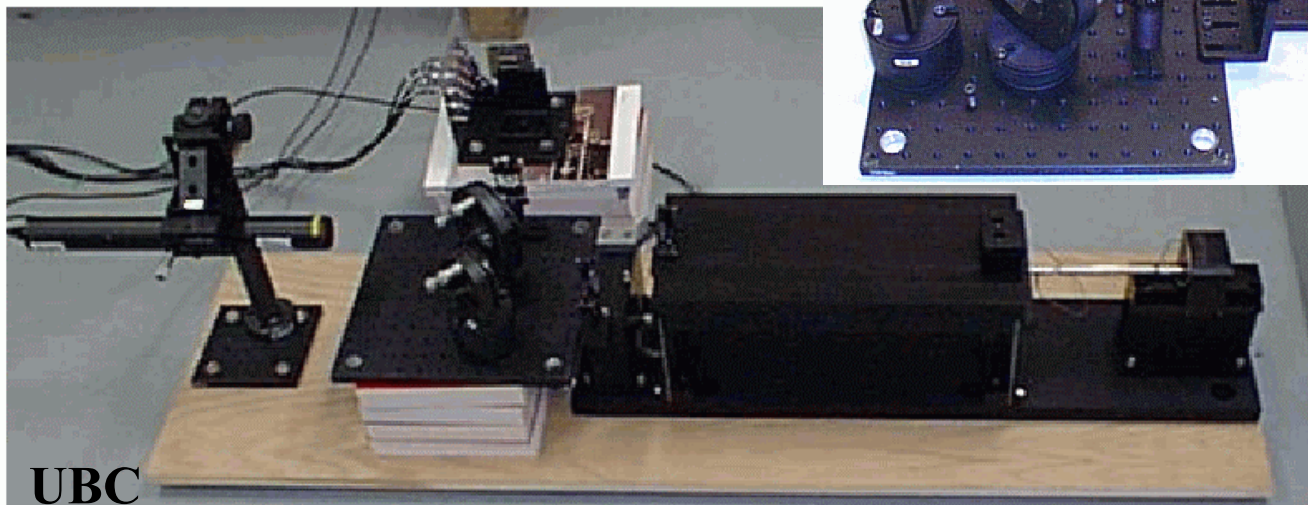


# R&D on mechanical stabilization with inertial and optical sensing



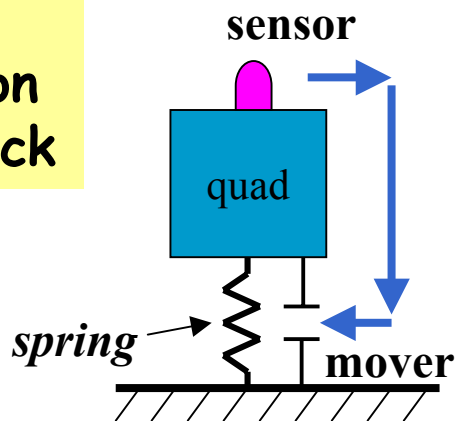
Talks of  
Joe Frisch  
Tom Mattison

and also  
Ralph Assmann  
for CLIC  
stabilization  
study



# Some questions on mechanical stabilization or field correction

## Position stabilization via feedback

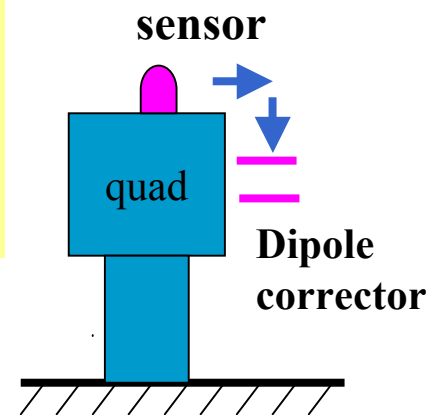


If FD is PM quad, how to deal with forces from the solenoid?

Estimated force on a PM quad can be 300 N to 2500 N, depending on configuration  
 [John Hodgson]  
 (The force is due to  $\mu > 1$  of PM material)

SC quad: talk of Brett Parker

## Correction of magnetic center via feedforward

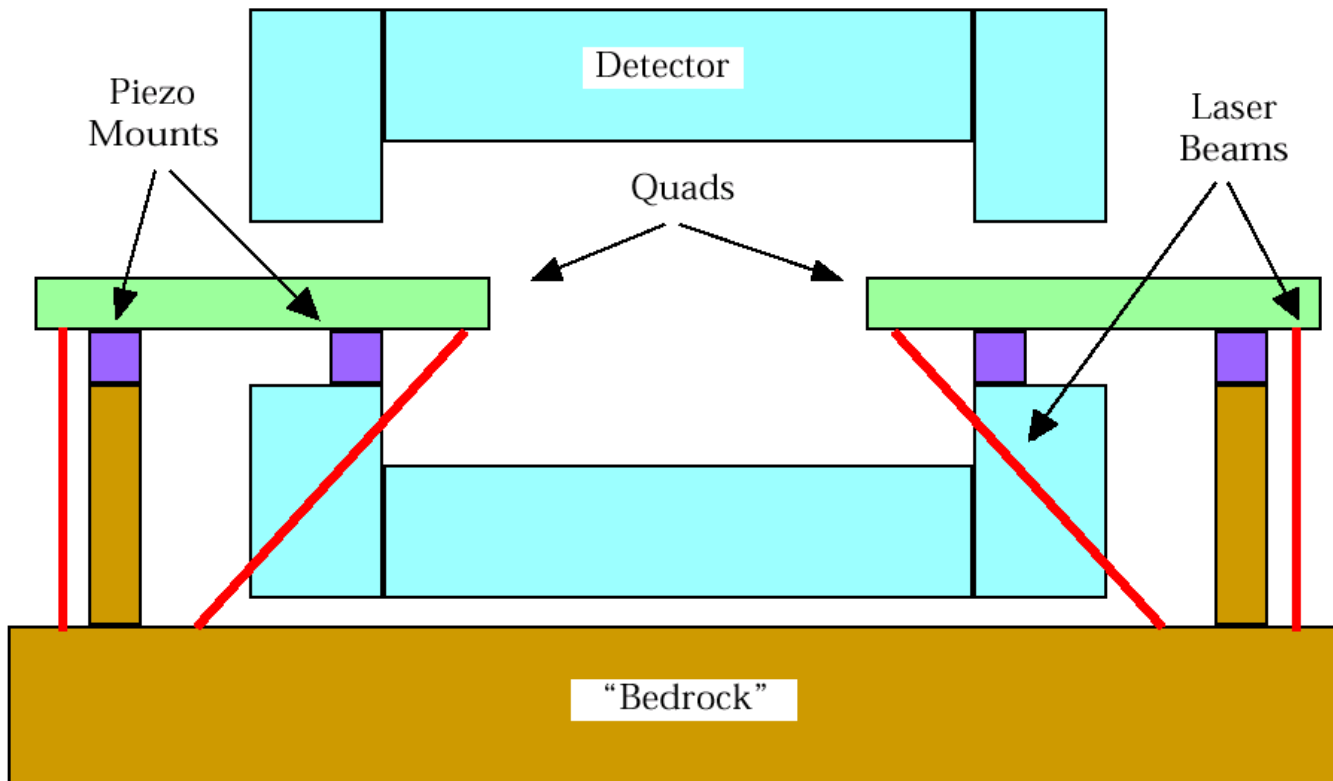


Possibly that much more vibration modes need to be controlled, more sensors, more complex algorithm?

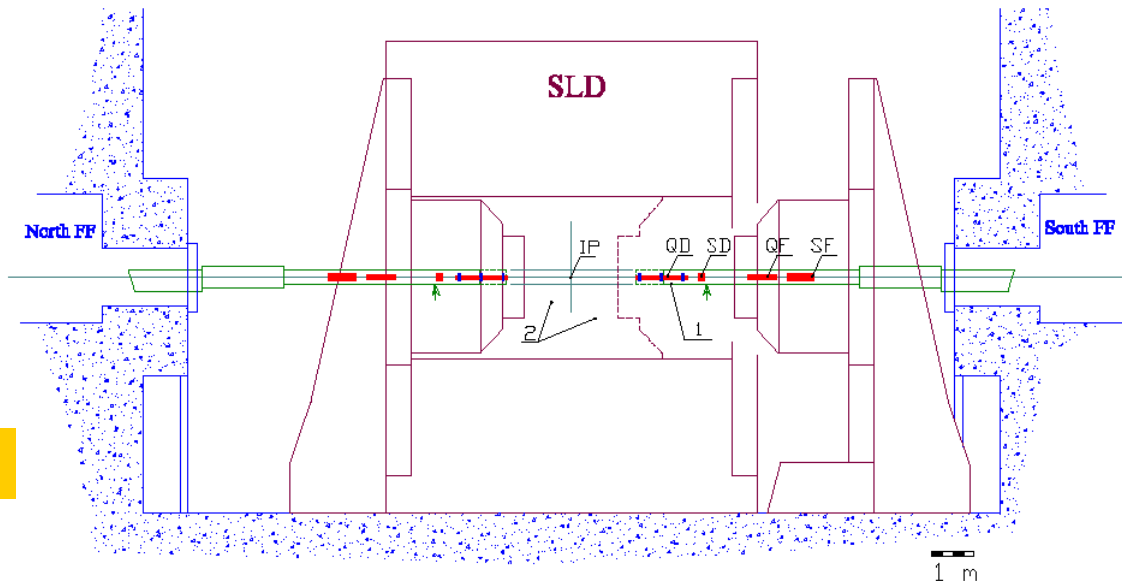
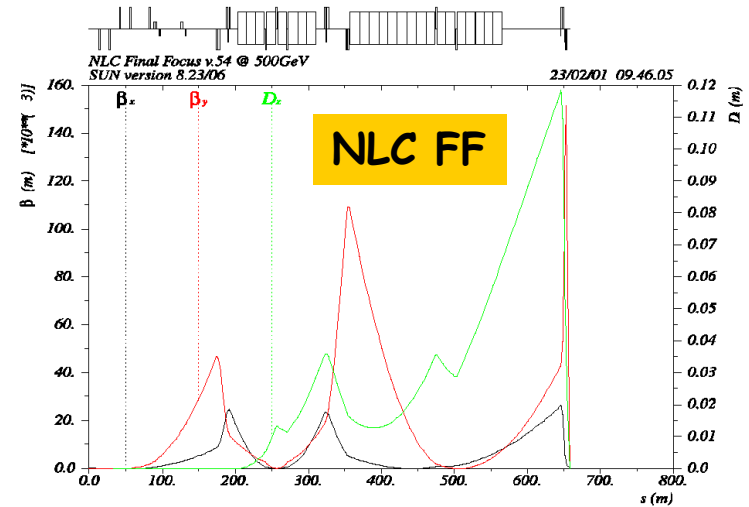
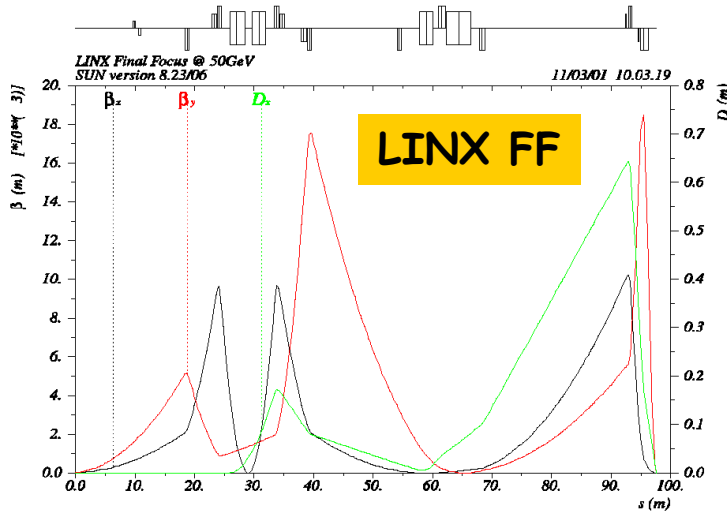
Less effective than feedback?

# Other questions to FD stability

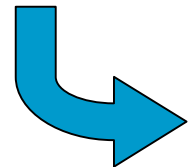
Do we support FD from noisier detector or only from tunnel, for the cost of much lower resonance frequency of the supporting girder? Other options?



# One of the goals of LINX facility is to master FD stabilization



LINX IR



Discussion on Tue PM and Thu AM, Talks of Tom Markiewicz, Mayda Velasco



# Joint efforts of many people

## SLAC

Chris Adolphsen  
 Fred Asiri  
 Gordon Bowden  
 Marty Breidenbach  
 John Cogan  
 Carlos Damian  
 Eric Doyle  
 Leif Eriksson  
 Joe Frisch  
 Linda Hendrickson  
 Tom Himel  
 Frederic Le Pimpec  
 Tom Markiewicz  
 Rainer Pitthan  
 Tor Raubenheimer  
 Robert Ruland  
 Andrei Seryi  
 Steve Smith  
 Peter Tenenbaum  
 Mike Woods  
 Nancy Yu

## CERN

Ralph Assmann  
 Stefano Redaelli

## FNAL

Joe Lach  
 Chris Laughton  
 Duane Plant  
 Vladimir Shiltsev

## BINP

Andrei Chupira  
 Anatoly Medvedko  
 Mikhail Kandaurov  
 Vasili Parkhomchuk  
 Shavkat Singatulin  
 Evgeny Shubin

## Northwestern Univ.

Heidi Shellman  
 Mayda Velasco

et al.

## UBC

Tom Mattison  
 Russ Greenall  
 Parry Fung

## Oxford

Phil Burrows  
 Simon Jolly  
 Gerald Myatt  
 Gavin Nesom  
 Colin Perry  
 Glen White

## Brookhaven

Nick Simos

## Stanford

Sri Adiga



## Summary

- There is good understanding of ground motion and vibration, and it is improving
  - But there may always be surprises
- There is a fair possibility that stability of LC luminosity can be provided
  - Provided that important issues are not left forgotten and are vigilantly pursued
- There are a lot of important details and particular concerns, that we should discuss during this Workshop